

Jet Propulsion Laboratory
California Institute of Technology

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JPL/NASA and Space Exploration

Insoo Jun

Jet Propulsion Laboratory, California Institute of Technology

November 6, 2018

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Topics

- Introduction:
 - JPL
 - Natural Space Environments (NSE) Group
 - Myself
- Solar System Exploration
 - Europa
 - Psyche
 - Mars Exploration Program
 - MSL
 - M2020

NASA Vision Statements



OUR MISSION

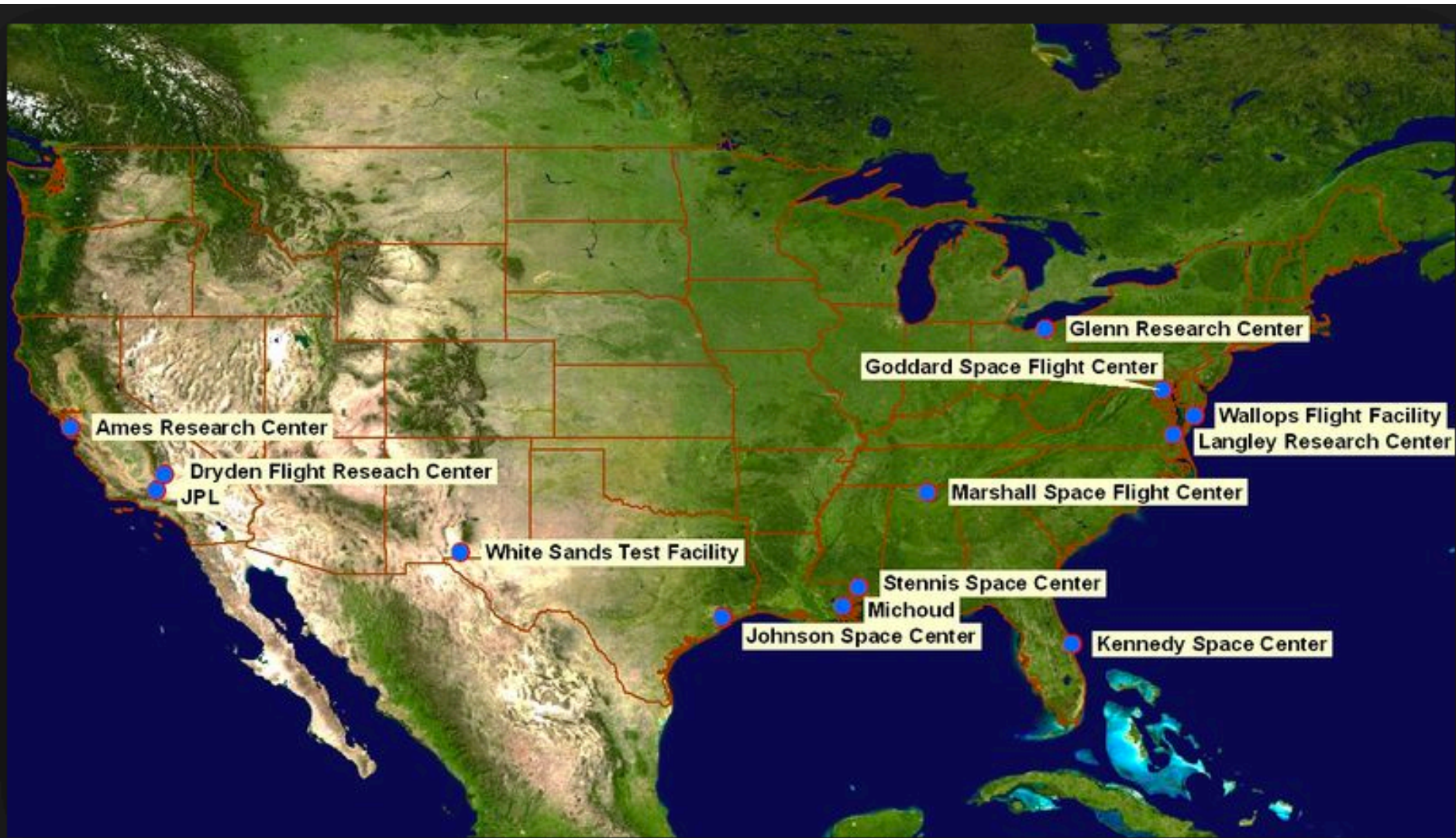
Drive advances in science, technology, aeronautics, and space exploration to enhance knowledge, education, innovation, economic vitality, and stewardship of Earth

Expand the frontiers of knowledge, capability, and opportunity in space

Serve the American public and accomplish our Mission by effectively managing our people, technical capabilities, and infrastructure

Advance understanding of Earth and develop technologies to improve the quality of life on our home planet

NASA Centers



Item	2017 Obama request	2017 Congress approved	2018 Trump request	2018 Omnibus
NASA TOTAL	\$19,025	\$19,653	\$19,092	\$20,736
SCIENCE				
Earth	\$2,032	\$1,921	\$1,754	\$1,921
Planetary	\$1,519	\$1,846	\$1,930	\$2,228
Astrophysics	\$782	\$750	\$817	\$850
JWST	\$569	\$569	\$534	\$534
Heliophysics	\$699	\$679	\$678	\$689
TOTAL	\$5,600	\$5,765	\$5,712	\$6,222
AERONAUTICS				
TOTAL	\$790	\$660	\$624	\$685
SPACE TECHNOLOGY				
TOTAL	\$827	\$687	\$679	\$760
EXPLORATION				
Orion	\$1,120	\$1,350	\$1,186	\$1,350
SLS	\$1,310	\$2,150	\$1,938	\$2,150
Ground systems	\$429	\$429	\$460	\$895
Exploration sys subtotal	\$2,859	\$3,929	\$3,584	\$4,395
R & D	\$477	\$395	\$350	\$395
TOTAL	\$3,337	\$4,324	\$3,934	\$4,790

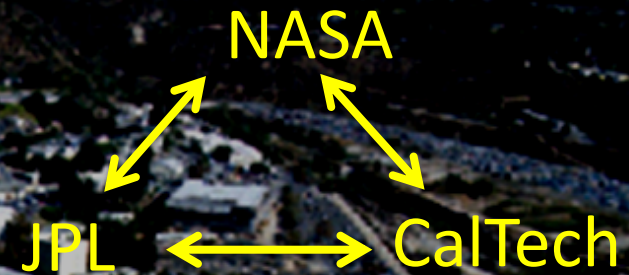
Science is the largest mission directorate within NASA

SPACE OPERATIONS				
Commercial crew	\$1,185		\$1,683	
Crew & cargo	\$1,573		\$732	
Space trans subtotal	\$2,758		\$2,415	
ISS	\$1,431		\$1,491	
TOTAL	\$5,076	\$4,951	\$4,741	\$4,752
EDUCATION				
TOTAL	\$100	\$100	\$37	\$100
SAFETY / SEC / SERVICES				
TOTAL	\$2,387	\$2,769	\$2,830	\$2,827
CONST / ENVIRO				
TOTAL	\$120	\$361	\$496	\$562
OIG				
TOTAL	\$38	\$38	\$39	\$39
NASA TOTAL	\$19,025	\$19,653	\$19,092	\$20,736

JPL is part of NASA and Caltech

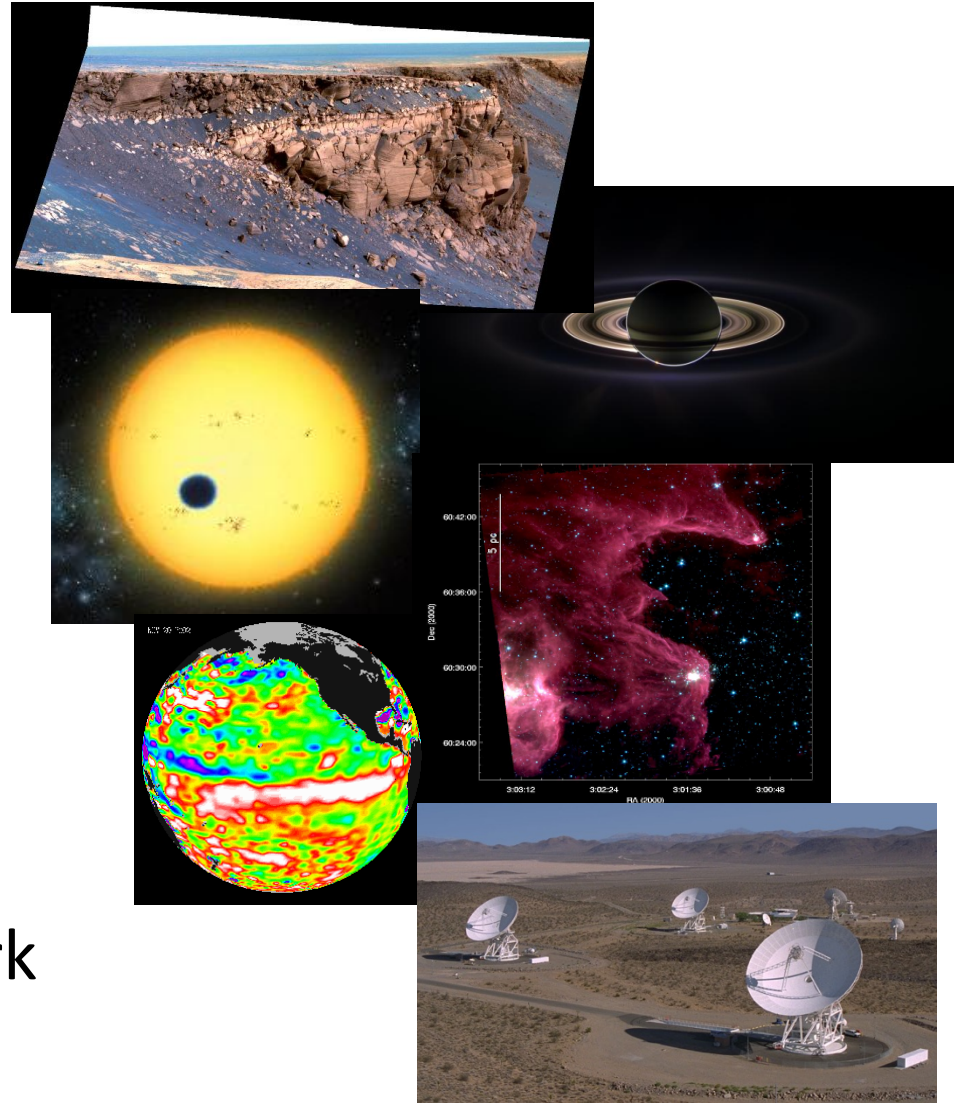


- Federally (NASA)-owned “Federally-Funded Research and Development Center” (FFRDC)
- University (Caltech)-operated
- >\$2 billion business base
- >6,000 employees
- 177 acres (Includes 22 acres leased for parking)
- 139 buildings and 36 trailers
- 673,000 net square feet of office space
- 906,000 net square feet of non-office space (e.g., labs)

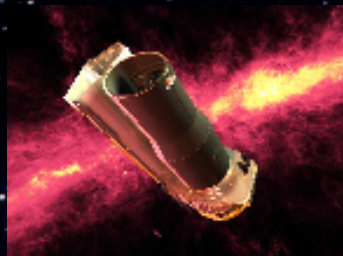


JPL's primary mission is robotic space exploration

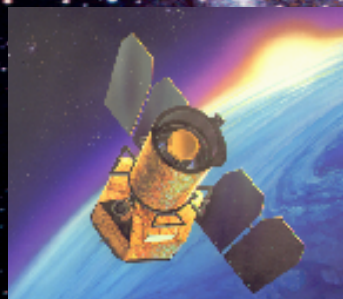
- Solar system
- Exoplanets
- Astrophysics
- Earth Science
- Interplanetary network



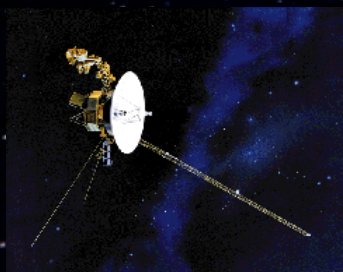
Thirty-nine spacecraft and instruments across the solar system (and beyond) – as of 2017



Spitzer



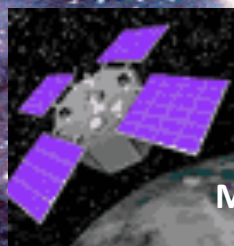
GALEX



Two Voyagers



Kepler



ACRIMSAT



Dawn



Wide-field Infrared Survey
Explorer (WISE)



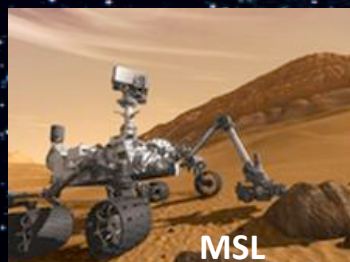
Mars Odyssey



Mars Reconnaissance Orbiter



Opportunity



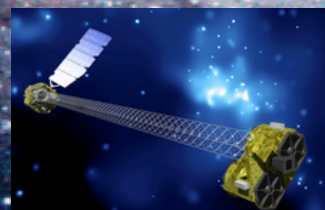
MSL



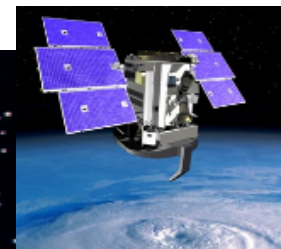
Cassini



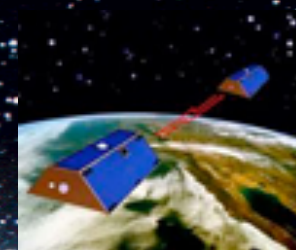
Juno



NuSTAR



CloudSat



GRACE

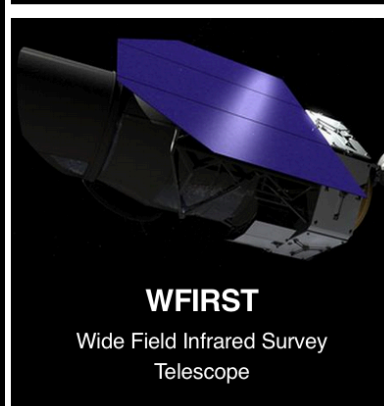
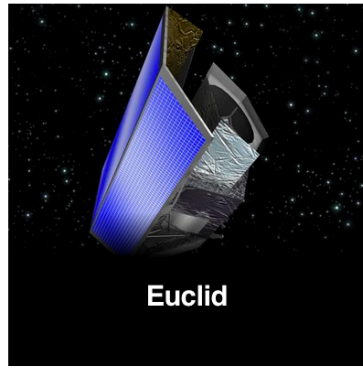
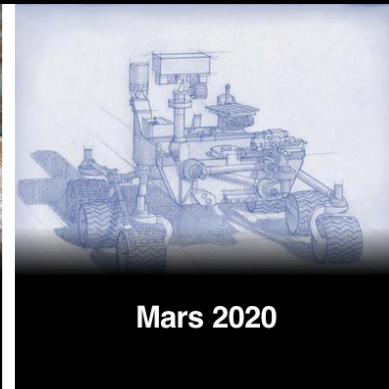
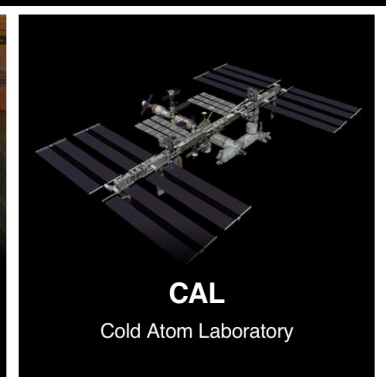
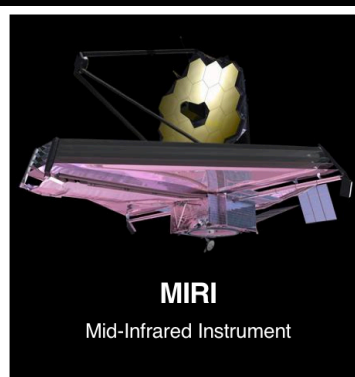
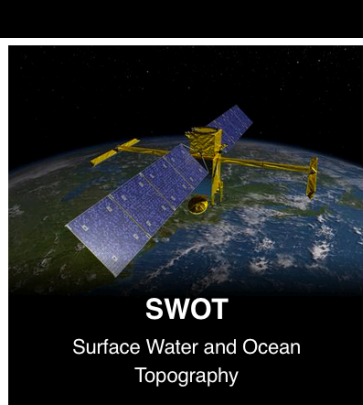


Jason 2 and Jason 3



SMAP

> 10 Future Missions (Planned)



Pre-Decisional Information
-- For Planning and
Discussion Purposes Only

JPL's Natural Space Environment Group

- Supports all JPL space-flight missions for the space environments and effects
 - Radiation (environment, shielding, charging,),
 - Meteoroids, Debris (MMOD)
 - Atomic Oxygen, Dust,
- Tasked with understanding environment hazards, to direct the design of robotic spacecraft we send into the solar system
- We strive to contribute to model development or improvement
 - e.g. models of solar proton fluence, outer-planet trapped radiation, meteoroids
 - Radiation environment monitoring
- Is the JPL lead for nuclear planetology (Gamma ray and neutron spectroscopy)

What do I do?

- Administrative
 - Group supervisor— Natural Space Environments Group, JPL
 - Chief Technologist for Reliability and Mission Environments Assurance Office
 - JPL lead for NASA Space Environment Capability Leadership Team (CLT)
- Scientific and Technical
 - Science Team:
 - Europa Clipper
 - MSL
 - Psyche
 - Europa Lander (Pre-project) – Radiation Lead
 - Co-lead for developing the Center for Space Radiation
 - Deputy Tech Fellow for for NASA Space Environment Technica
 - Space Weather Architecture WG
- Thesis advisor for two PhD students (MIT and TAMU)
 - MIT done
 - Looking for another one for Juno data analysis
- On-going collaboration with Universities
 - MDAP (with University of Tennessee) – MSL DAN and RAD dat
 - SURP (with University of Colorado) – Juno data analysis

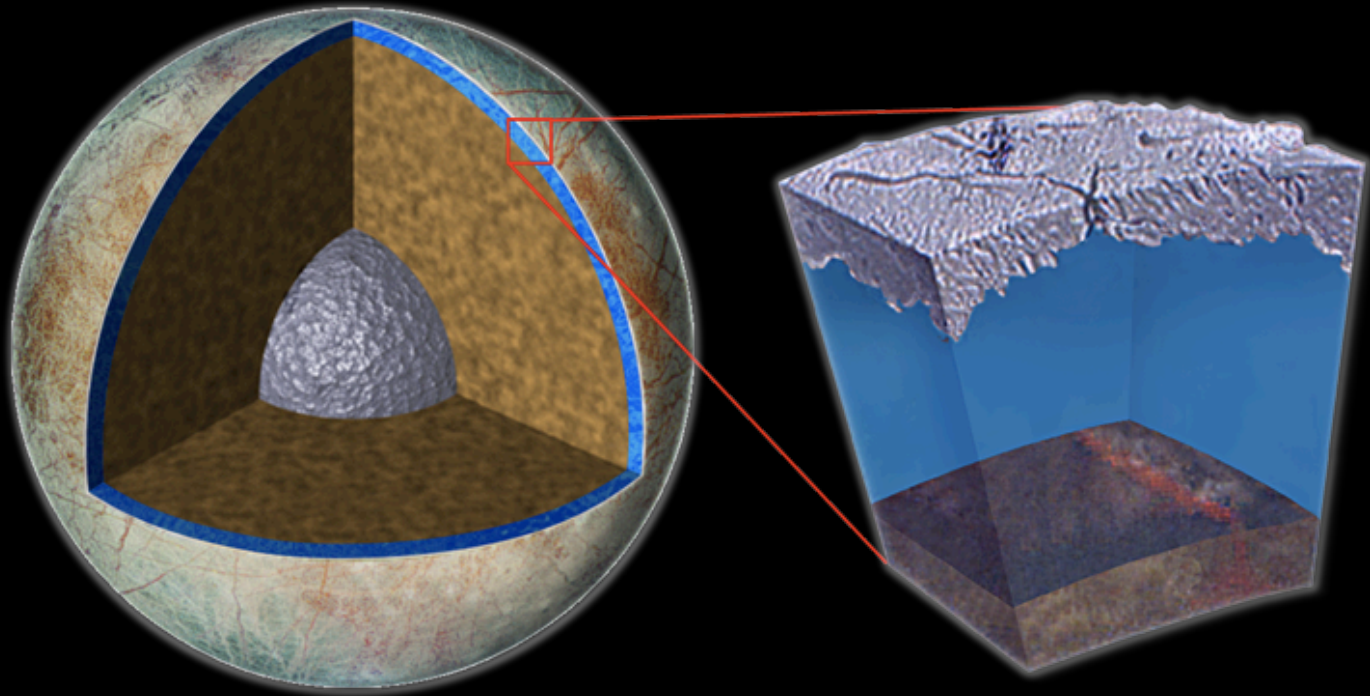


Recent Recognition

- 2018 NASA Exceptional Public Service Medal (EPSM)
 - *For exceptional service in the field of space radiation, developing engineering capabilities and mitigating damages to enable NASA mission success*
- Citation to the Europa Clipper Science Team
 - *Dr. Jun is the Chief Technologist He is a world expert in the radiation environment of Europa and the Jupiter system, and he has worked extensively on all of JPL's Jupiter mission projects and studies for radiation-related issues, typically as the radiation lead.*

Europa Mission

The Ocean That Beckons



*Europa, with its probable vast subsurface ocean sandwiched between a potentially active silicate interior and a highly dynamic surface ice shell, offers **one of the most promising extraterrestrial habitable environments**, and a plausible model for habitable environments beyond our solar system*

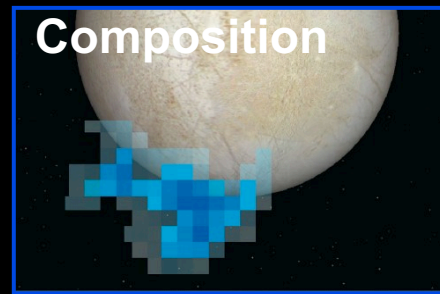
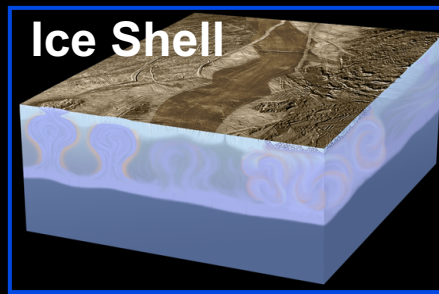
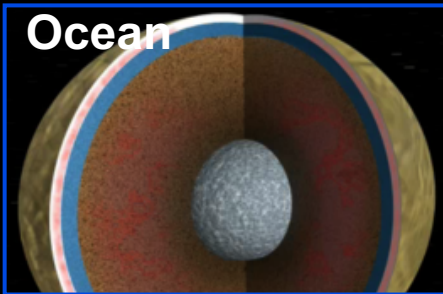
The Planetary Decadal Survey, 2011

Achieving Decadal Science

“The first step in understanding the potential of the outer solar system as an abode for life is a Europa mission with the goal of

- Confirming the presence of an interior ocean,*
- Characterizing the satellite’s ice shell, and*
- Enabling understanding of its geologic history”*

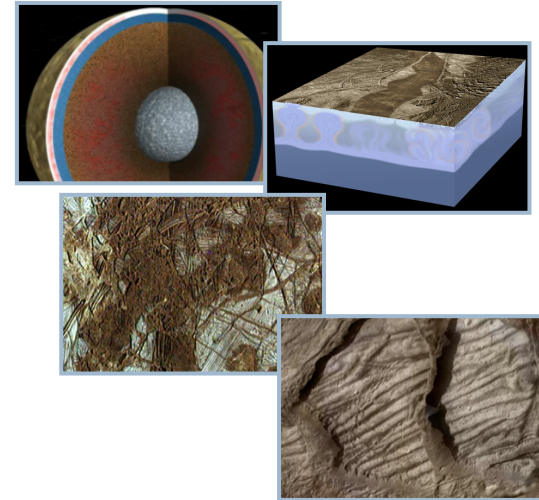
– The Planetary Decadal Survey, 2011



Europa Clipper Goals and Objectives

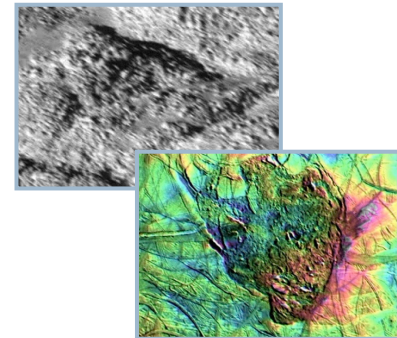
- **Science Goal:** *Explore Europa to investigate its habitability*

- **Ocean & Ice Shell:** Characterize ice shell and subsurface water, including heterogeneity, ocean properties, and surface-ice-ocean exchange
- **Composition:** Understand habitability of Europa's ocean through composition and chemistry
- **Geology:** Understand formation of surface features, including sites of recent or current activity, and characterize high science interest localities



- **Reconnaissance Goal:** *Characterize Safe and Scientifically Compelling Sites for a Future Lander Mission to Europa*

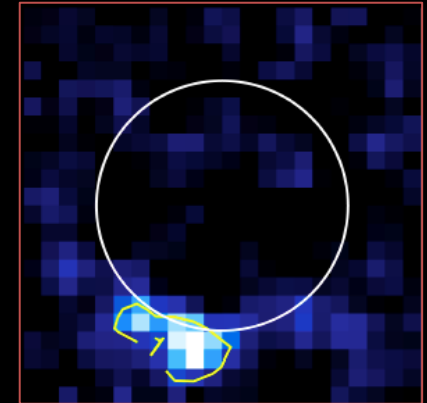
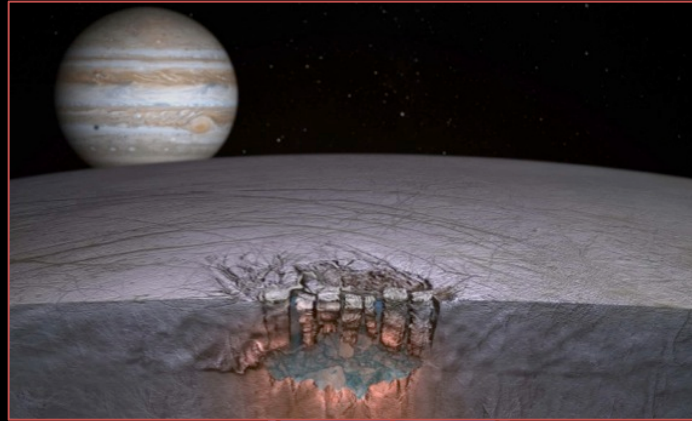
- Distribution of surface hazards, load-bearing capacity of surface, structure of the subsurface, and regolith thickness
- Composition of surface materials, geologic context, potential for geologic activity, proximity of near surface water, and potential for active upwelling of ocean material



Europa: Ingredients for Life?



Water: Are a global ocean and lakes hidden by Europa's shell of ice?

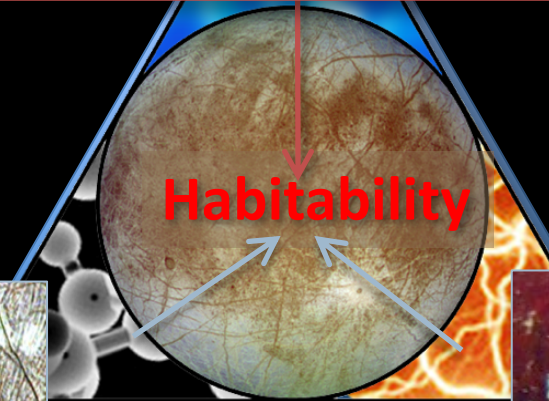


Plumes at Europa!
(Roth et al., 2013)

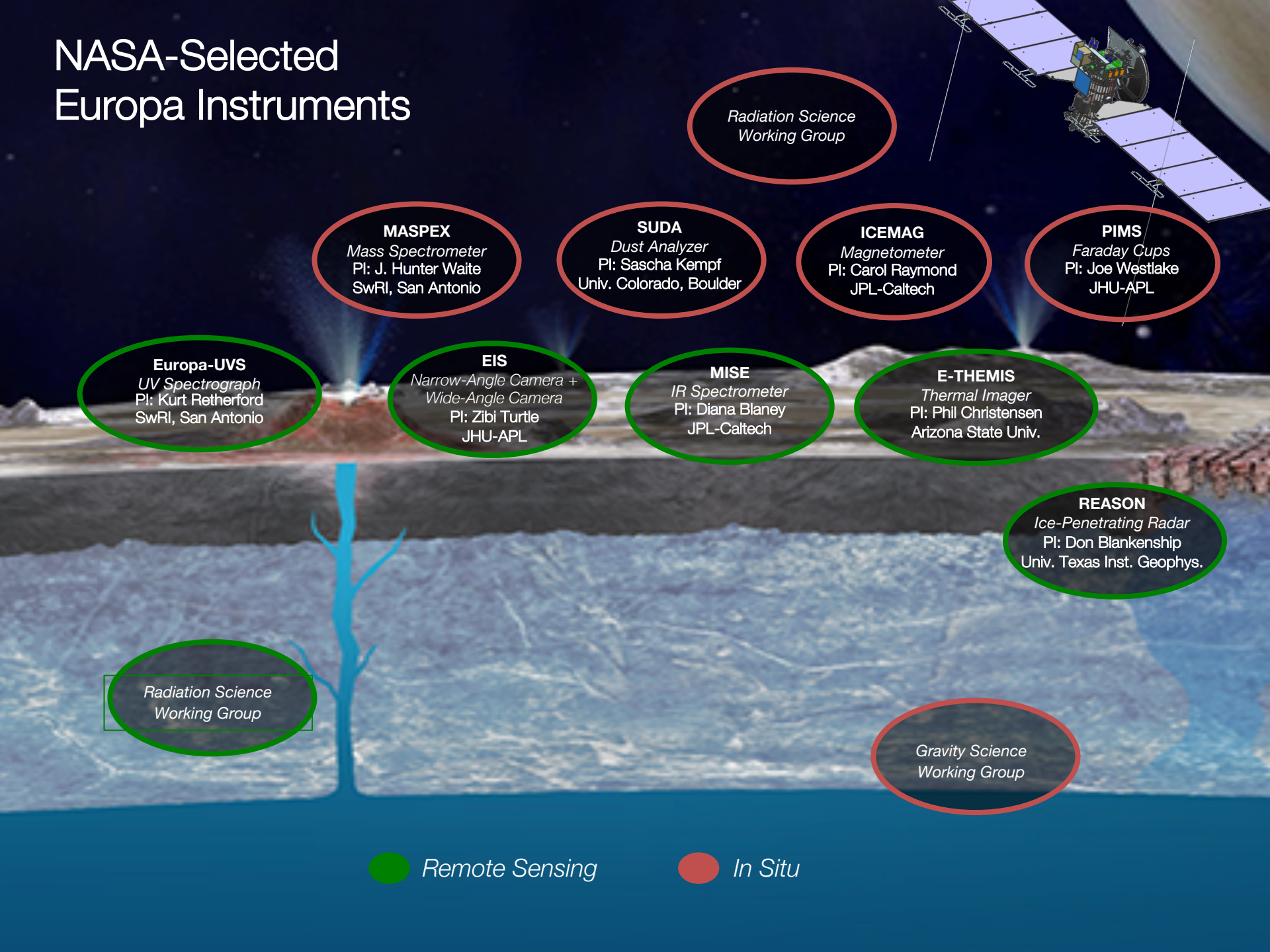
Chemistry: Do red surface deposits contain organics from below?



Energy: Can surface oxidants provide energy for metabolism?



NASA-Selected Europa Instruments



Radiation Science
Working Group

MASPEX

Mass Spectrometer
PI: J. Hunter Waite
SwRI, San Antonio

SUDA

Dust Analyzer
PI: Sascha Kempf
Univ. Colorado, Boulder

ICEMAG

Magnetometer
PI: Carol Raymond
JPL-Caltech

PIMS

Faraday Cups
PI: Joe Westlake
JHU-APL

Europa-UVS

UV Spectrograph
PI: Kurt Retherford
SwRI, San Antonio

EIS

Narrow-Angle Camera +
Wide-Angle Camera
PI: Zibi Turtle
JHU-APL

MISE

IR Spectrometer
PI: Diana Blaney
JPL-Caltech

E-THEMIS

Thermal Imager
PI: Phil Christensen
Arizona State Univ.

REASON

Ice-Penetrating Radar
PI: Don Blankenship
Univ. Texas Inst. Geophys.

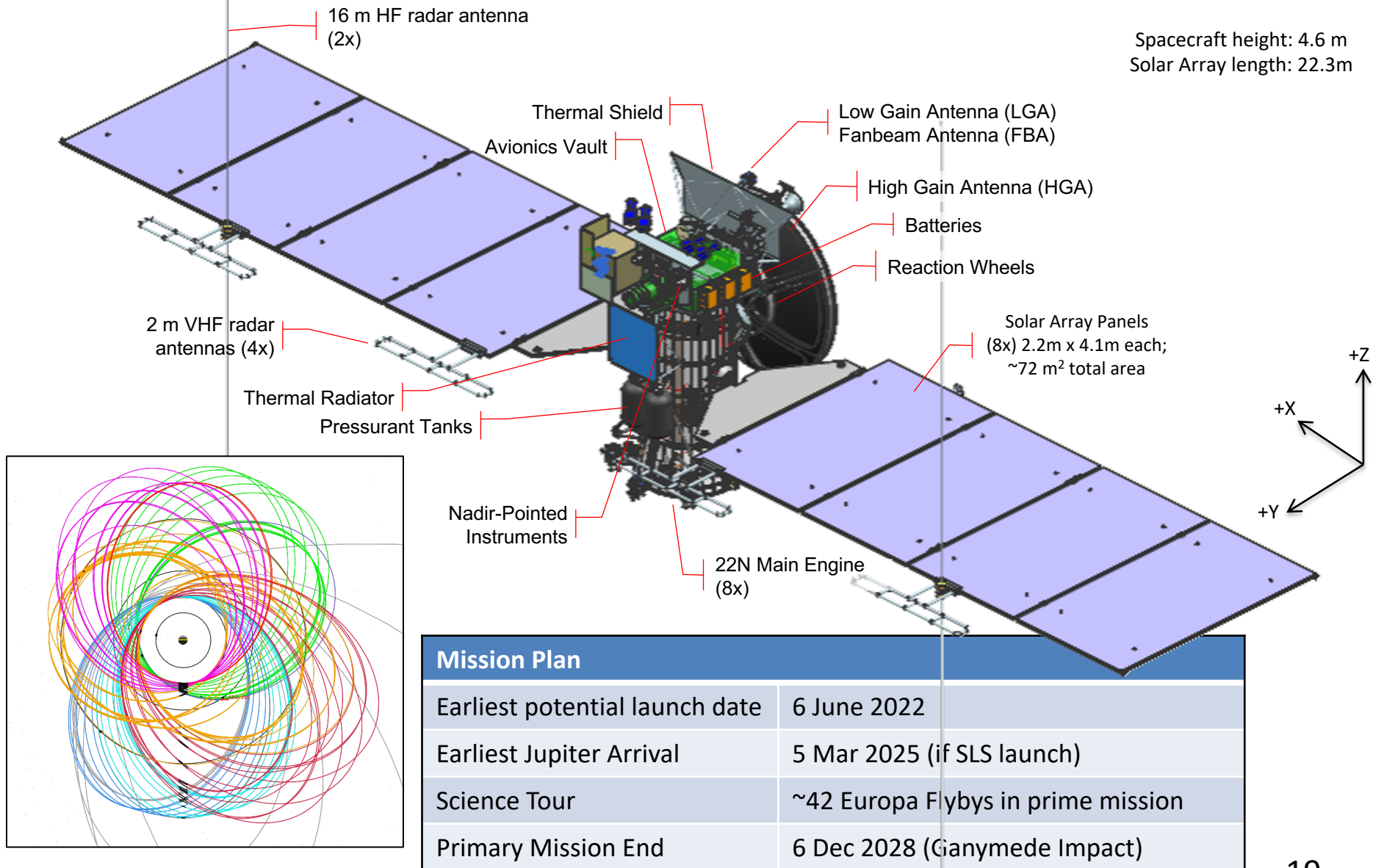
Radiation Science
Working Group

Gravity Science
Working Group

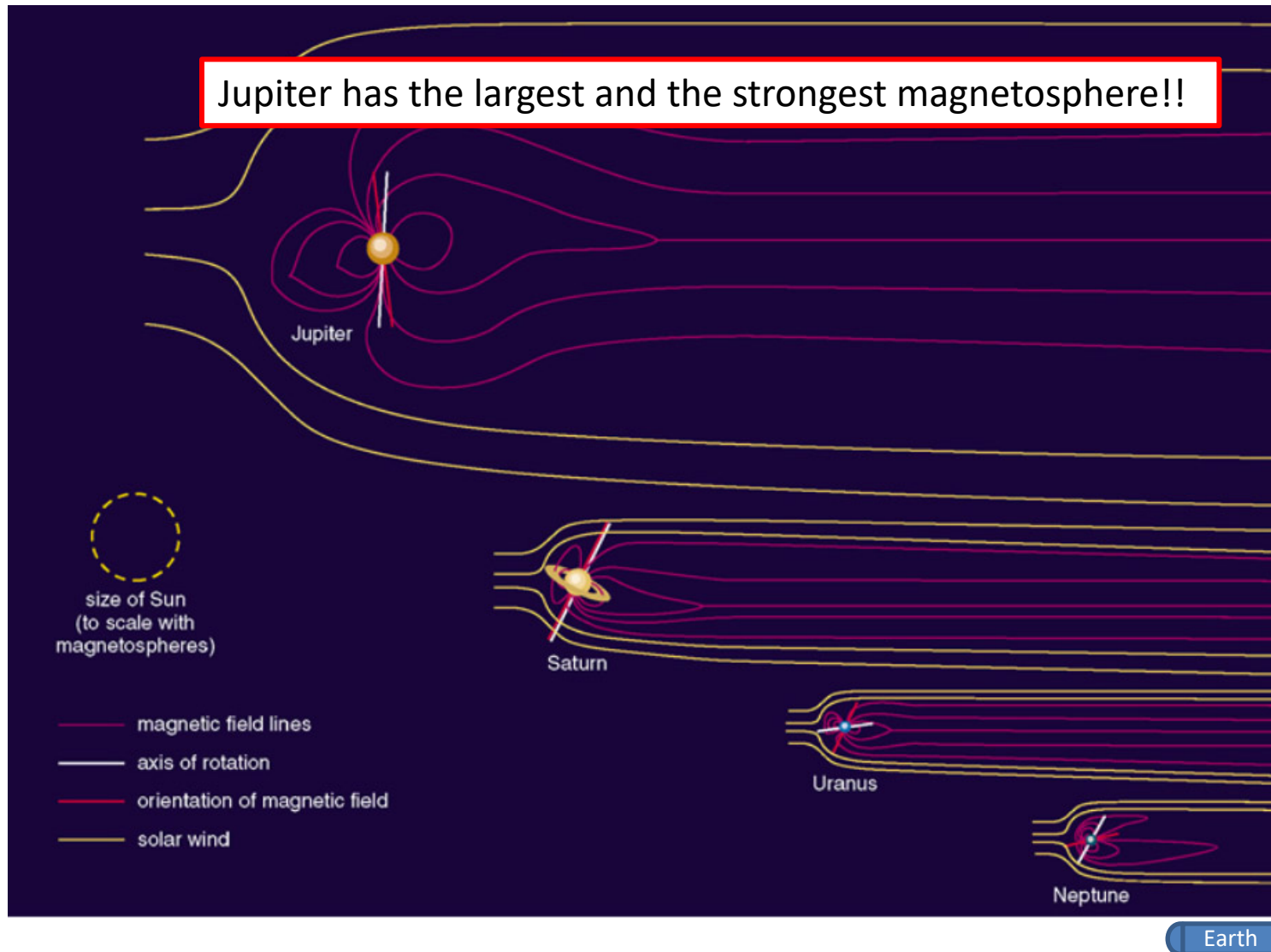
Remote Sensing

In Situ

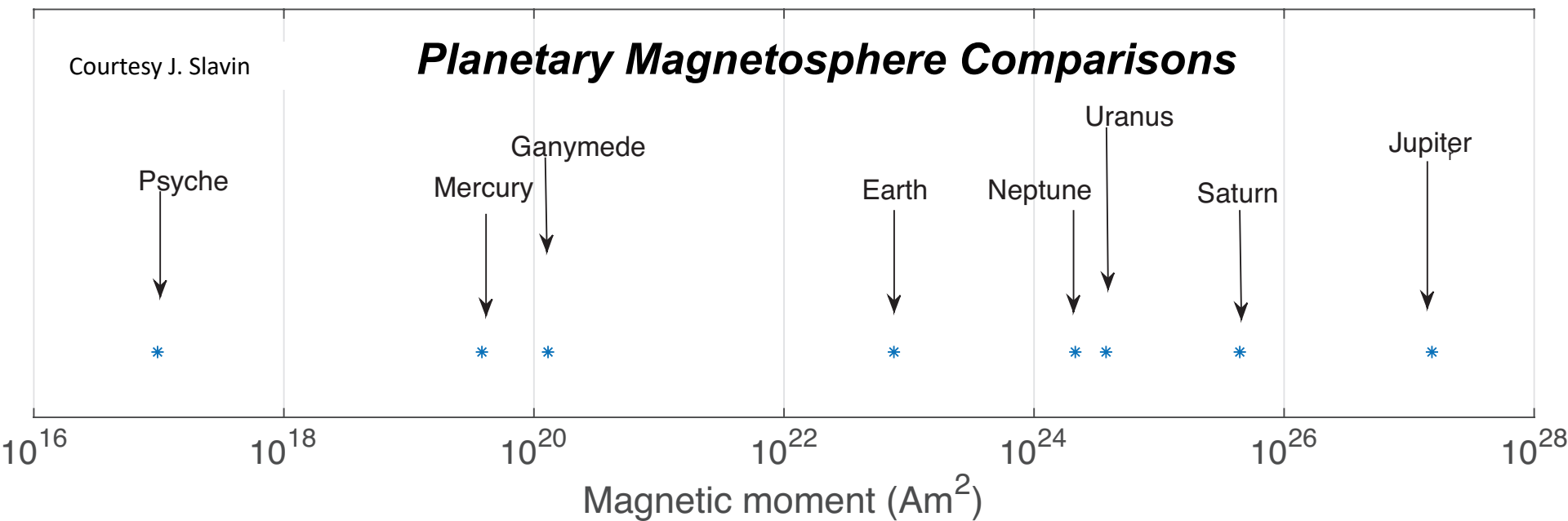
Europa Mission Concept Overview



Planetary Magnetospheres in the Solar System



Planetary Characteristics



PHYSICAL PROPERTIES:	Earth	Jupiter	Saturn	Uranus	Neptune
Equatorial radius (km)	6378	71492	60268	25559	24766
Mass (kg)	5.97E+24	1.90E+27	5.68E+26	8.68E+25	1.02E+26
Semi-major axis (AU)	1	5.2	9.54	19.19	30.07
Sidereal day (hr)	23.93	9.89	10.61	17.14	16.7
Dipole tilt (deg)	11.3	9.6	0	58.6	47
Dipole offset (rp)	0.0725	0.131	0.04	0.3	0.55
Magnetic moment (G-Rp ³)	0.305	4.28	0.21	0.228	0.133

Jupiter's Magnetosphere

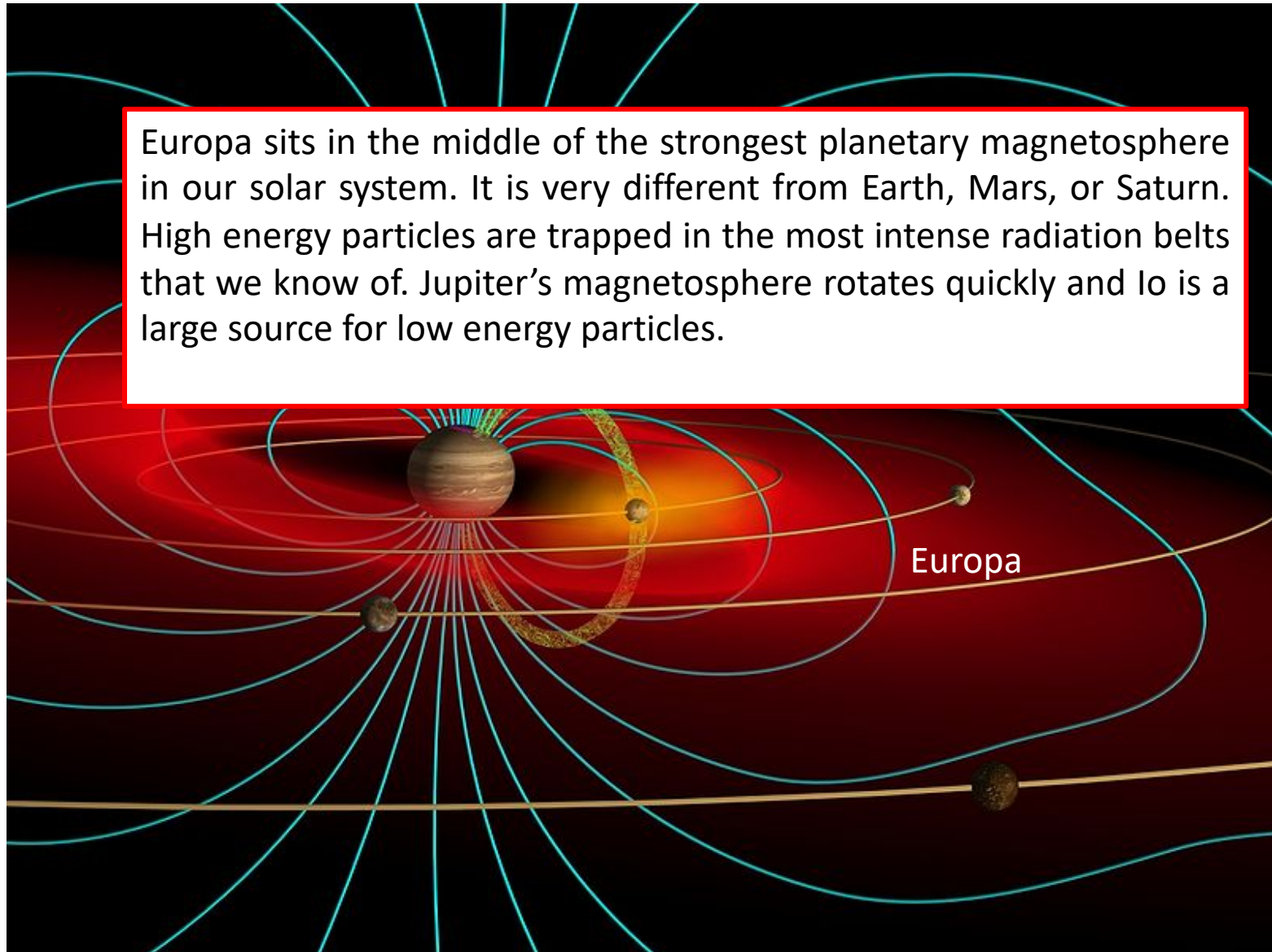
	Earth	Jupiter
Equatorial radius (km)	6.38×10^3	7.14×10^4
Magnetic moment (G-cm ³)	8.1×10^{25}	1.59×10^{30}
Rotation period (hr)	24.0	10.0
Aphelion/perihelion (AU)	1.01/0.98	5.45/4.95

- Jupiter is roughly 10 times the size of the Earth while its magnetic moment is 20,000 times larger.
- As the magnetic field at the equator is proportional to the magnetic moment divided by the cube of the radial distance, the Jovian magnetic field is proportionally **20 times** larger than the Earth's.

The energy and flux levels of trapped particles in the Jovian system can be much higher than those at the Earth or in the interplanetary space

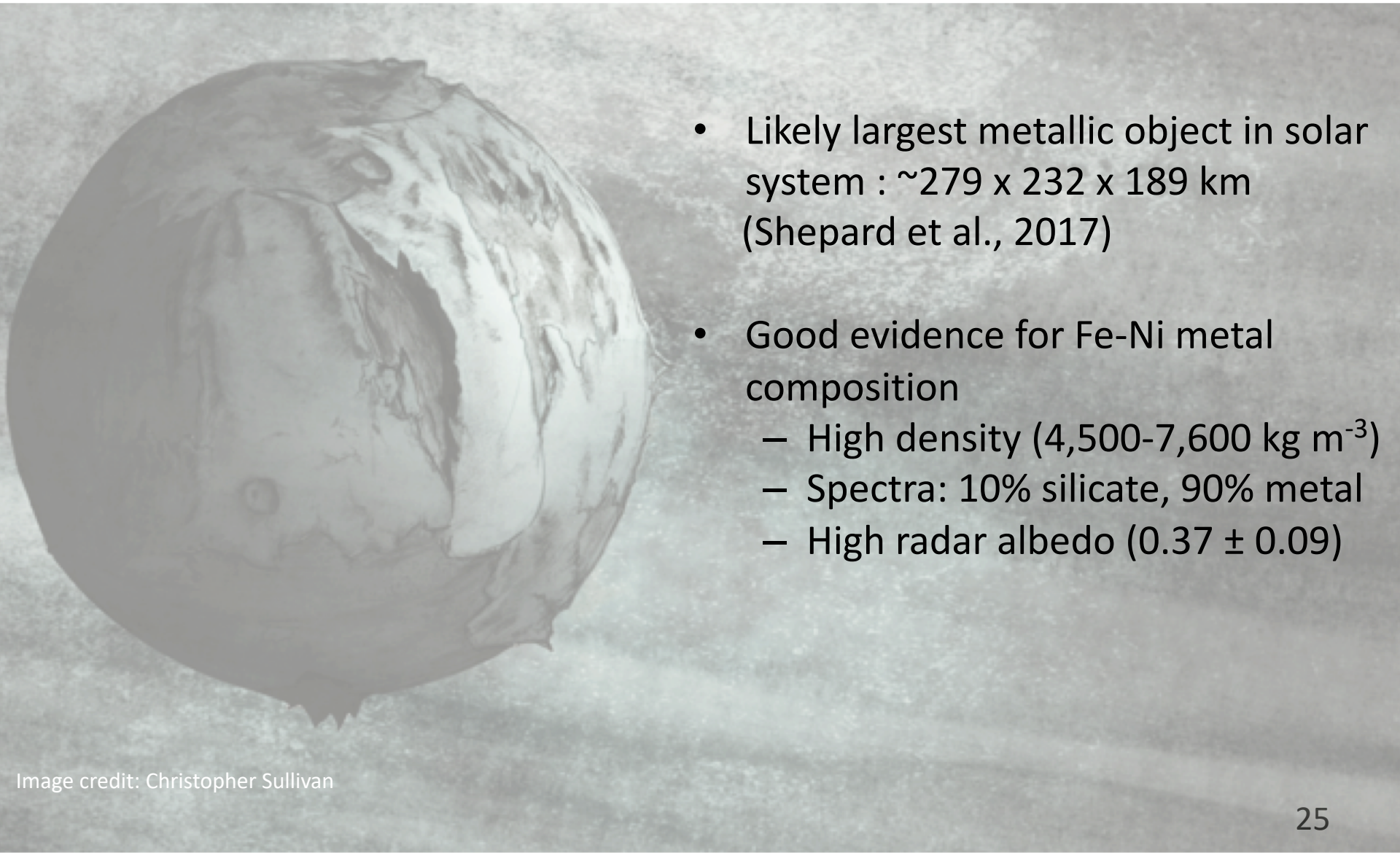
Jupiter

Europa sits in the middle of the strongest planetary magnetosphere in our solar system. It is very different from Earth, Mars, or Saturn. High energy particles are trapped in the most intense radiation belts that we know of. Jupiter's magnetosphere rotates quickly and Io is a large source for low energy particles.



Psyche Mission

Why (16) Psyche?



- Likely largest metallic object in solar system : $\sim 279 \times 232 \times 189$ km (Shepard et al., 2017)
- Good evidence for Fe-Ni metal composition
 - High density ($4,500\text{--}7,600$ kg m⁻³)
 - Spectra: 10% silicate, 90% metal
 - High radar albedo (0.37 ± 0.09)

Psyche: Looking inside planets with a journey to a metal world

Science Goals

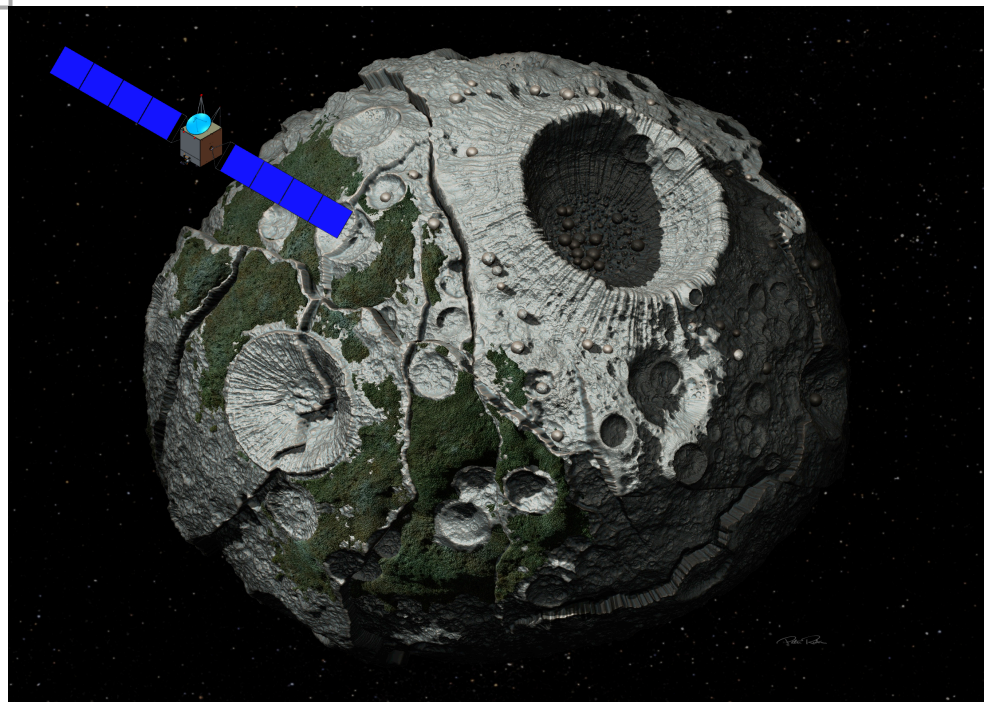
- **Look back in time.** Understand a previously unexplored component of the early building blocks of planets: *iron cores*.
- **Look inside the terrestrial planets**, including Earth, by directly examining the interior of a differentiated body, which otherwise could not be seen.
- **Explore a new type of world.** For the first time, examine a world made not of rock or ice, but of iron.

Mission Description

- Target: (16) Psyche
- Spacecraft Partner: JPL flight system w/SSL SEP chassis
- Launch Date: Nov 2021
- Arrival Date: Jan 2026
- Science Phase duration: 12 months

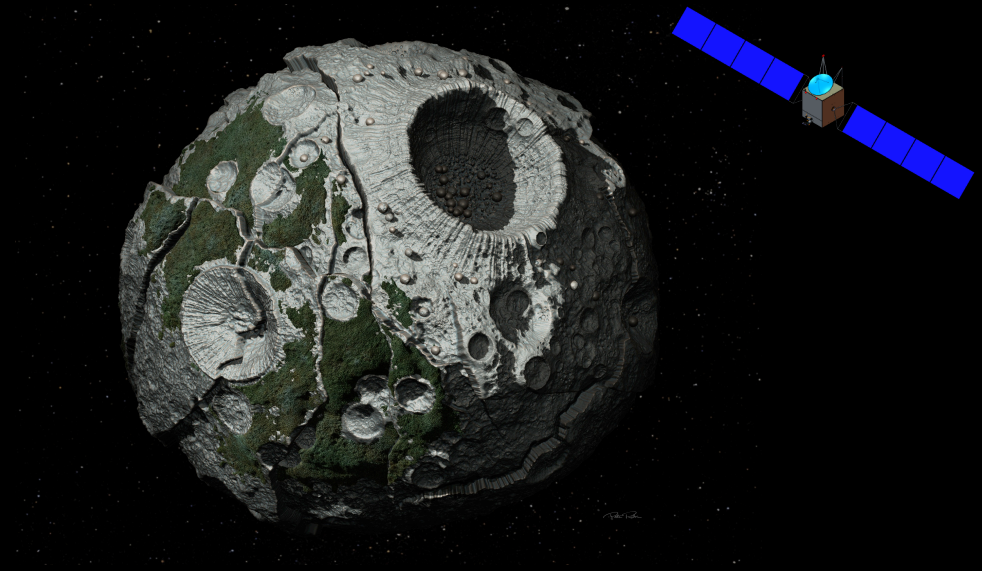
Payload

- Magnetometer: MIT/UCLA (MMS/Insight)
- Multispectral imager x 2: ASU/MSSS (MSL heritage)
- **Gamma Ray/Neutron Spectrometers: APL (MESSENGER heritage)**
- Gravity Science: X-band telecom, no dedicated hardware

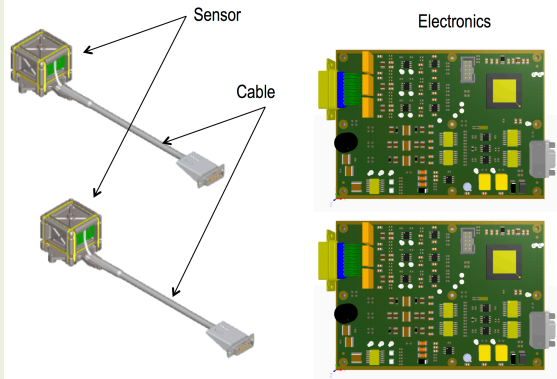


Science Objectives

- A. Determine whether Psyche is a core, or if it is unmelted material.
- B. Determine the epoch of formation of Psyche by measuring the relative ages of regions of its surface.
- C. Determine whether small metal bodies incorporate the same light elements as are expected in the Earth's high-pressure core.
- D. Determine whether Psyche was formed under more oxidizing or more reducing conditions than Earth's core.
- E. Characterize Psyche's topography.

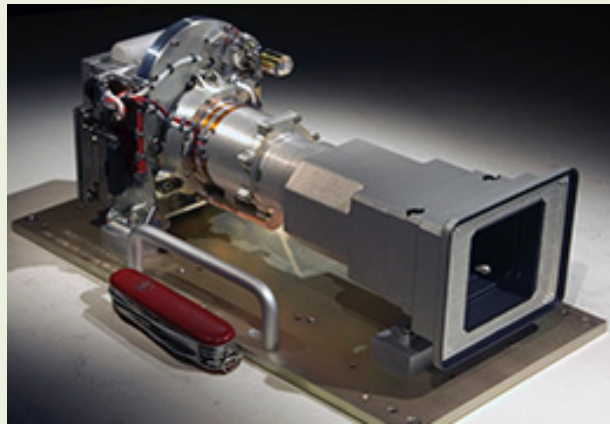


Psyche's Instrument Suite



Magnetometer

- MIT lead/UCLA built
- 2 sensors
- 0.66 kg total
- 2-m boom (gradiometer)
- 0.2 – 100,000 nT range
- MMS, Insight heritage

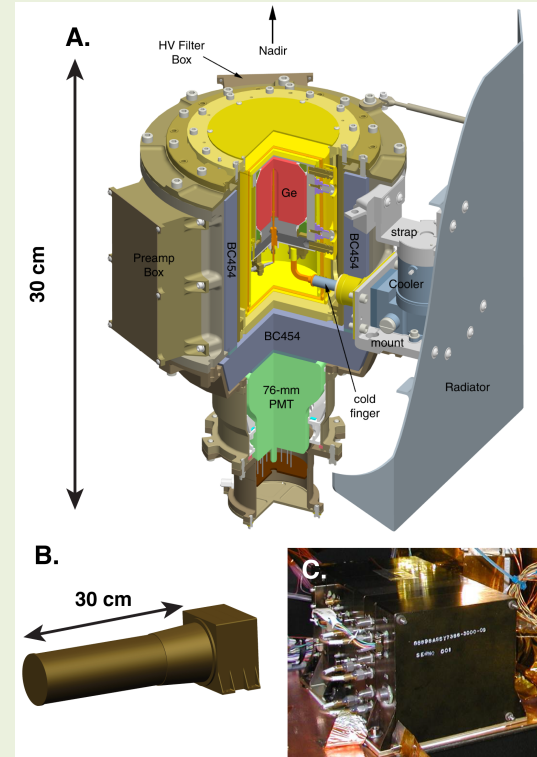


Multispectral Imager

- ASU lead/MSSS built
- Redundant units
- 2.9 kg each
- 8 filters
- 35 m/pixel highest orbit
- 5 m/pixel lowest orbit
- MSL Mastcam heritage

Gamma Ray and Neutron Spectrometer

- APL
- 9.6 kg, 2-m boom
- High Purity Ge detector
- He³ sensor (thermal neutrons)
- MESSENGER heritage



THANK YOU!

QUESTIONS?

BACKUP

NASA'S JOURNEY TO

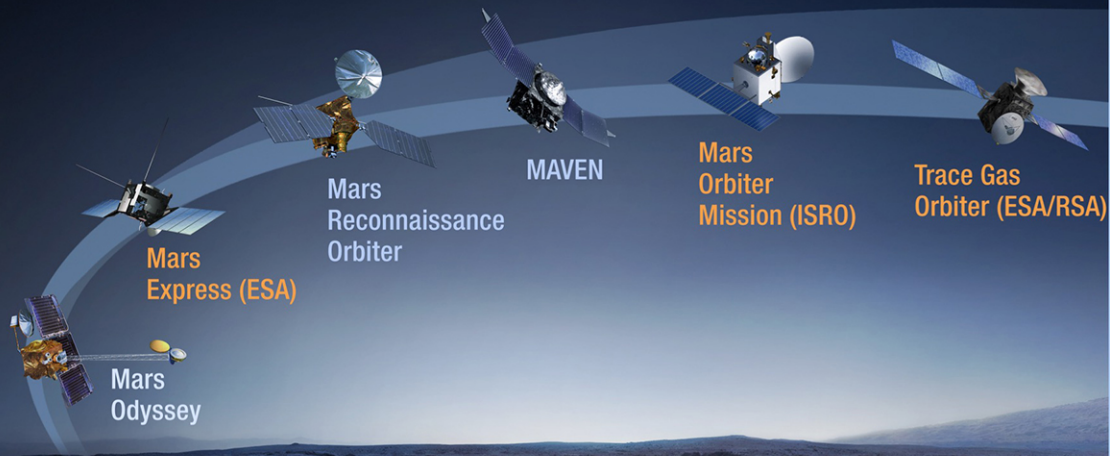
MARS



Operational 2001–2017

2018

2020 and Beyond



Planning for the Future

ExoMars Rover (ESA/RSA)

SpaceX Dragon

InSight

M2020 Rover

Mars Rover (China)

Explore Habitability

Seek Signs of Life

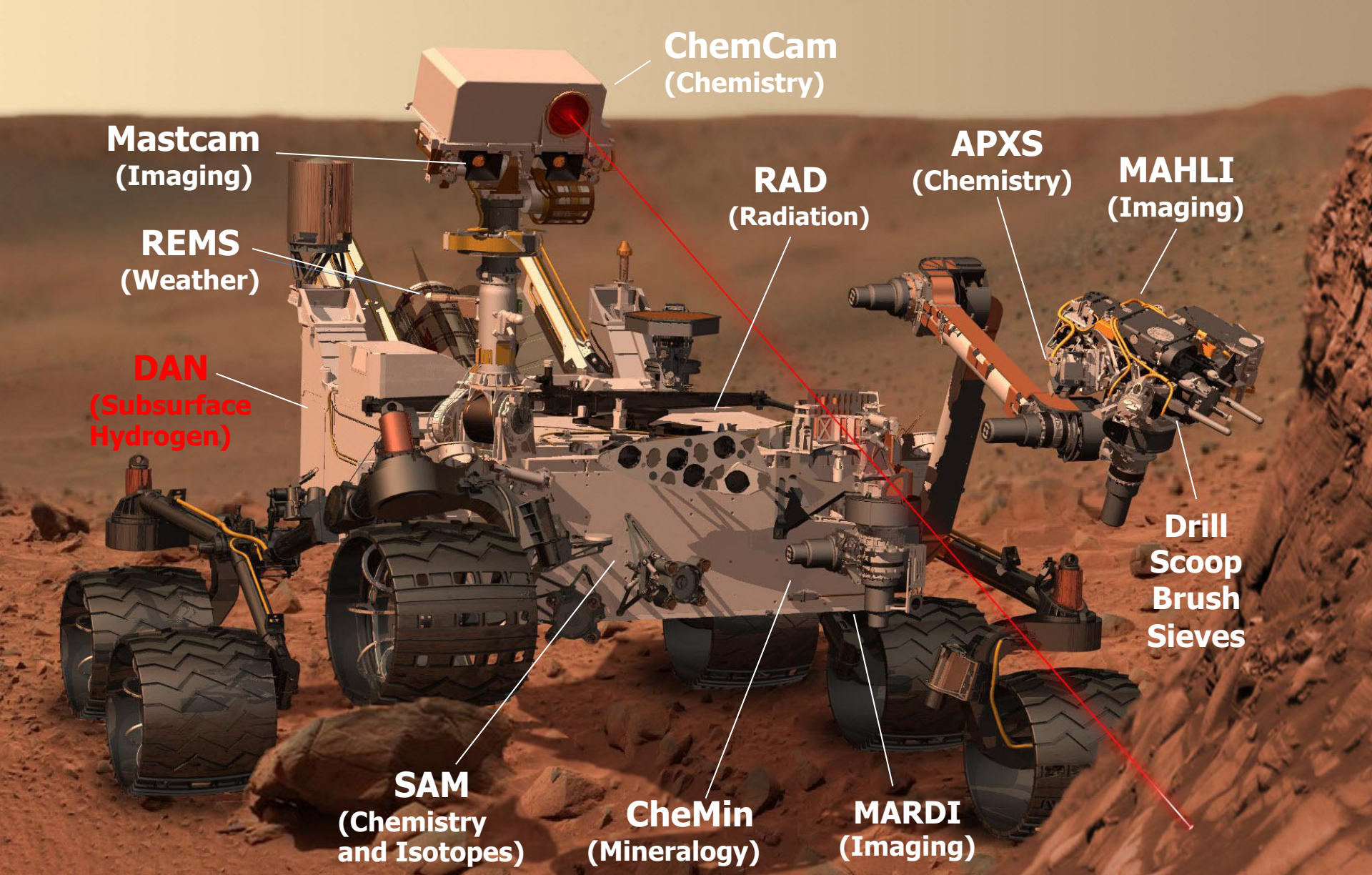
Prepare for Future Human Explorers

MARS SCIENCE LABORATORY (CURIOSITY)

Curiosity's primary scientific goal is to explore and quantitatively assess a local region on Mars' surface as a potential habitat for life, past or present

- Biological potential
- Geology and geochemistry
- Water and weather
- Radiation hazards

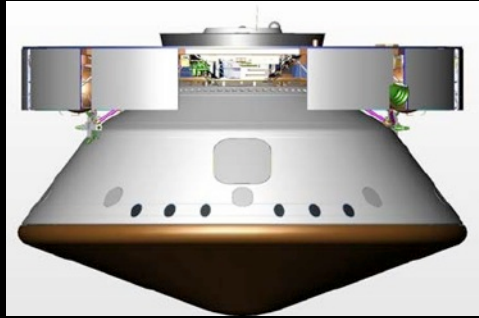




Curiosity's Science Payload



Landing and Mobility



CRUISE/APPROACH

- 8-month cruise
- Arrived August 5th, 2012

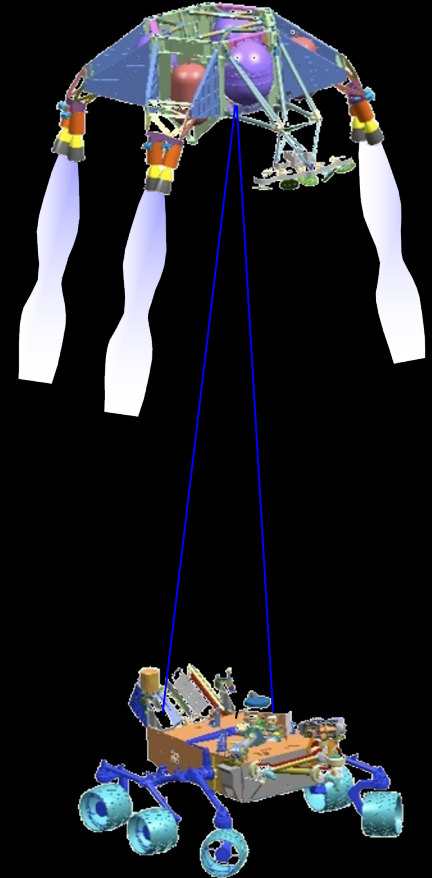
LAUNCH

- Nov-2011
- Atlas V (541)



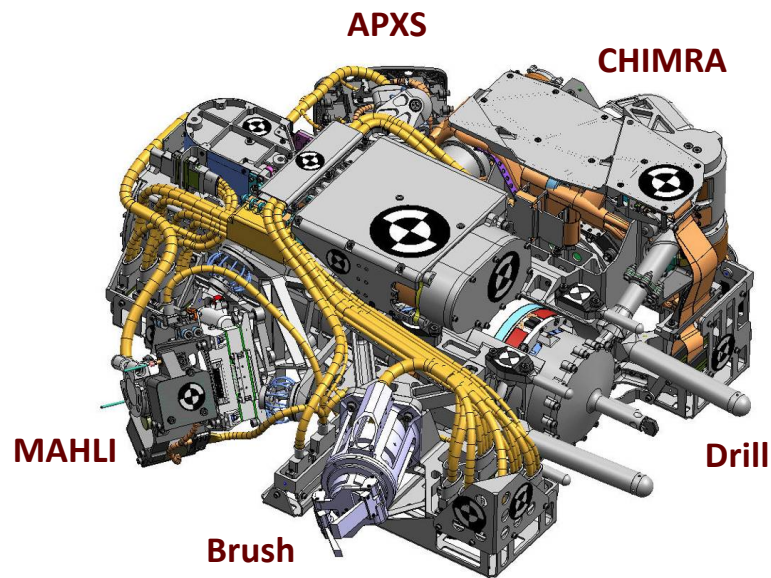
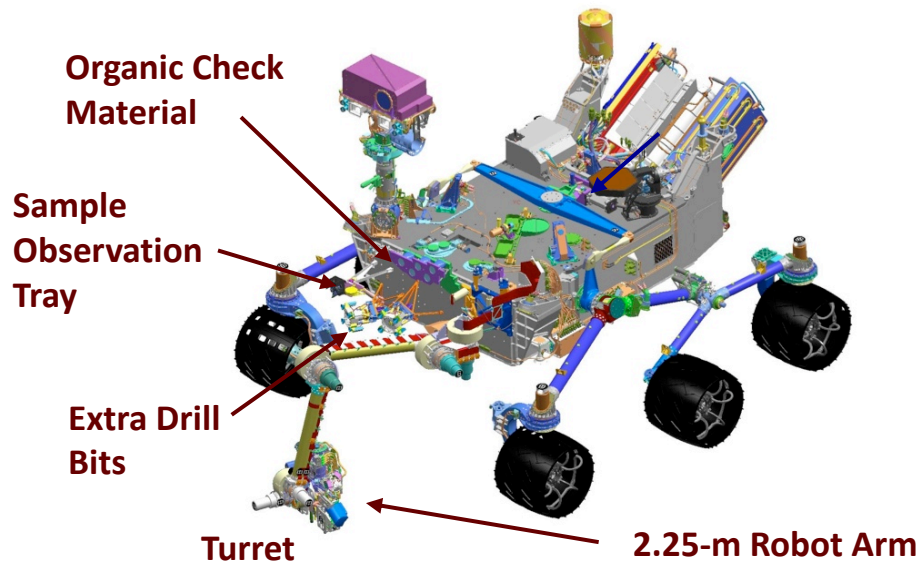
ENTRY, DESCENT, LANDING

- Guided entry and powered “sky crane” descent
- 20×25-km landing ellipse
- Access to landing sites $\pm 30^\circ$ latitude, < 0 km elevation
- 900 kg rover



SURFACE MISSION

- Long-lived Plutonium power source
- Ability to drive out of landing ellipse, up 100 m per sol
- 84 kg of science payload
- Direct (uplink) and relayed (downlink) communication



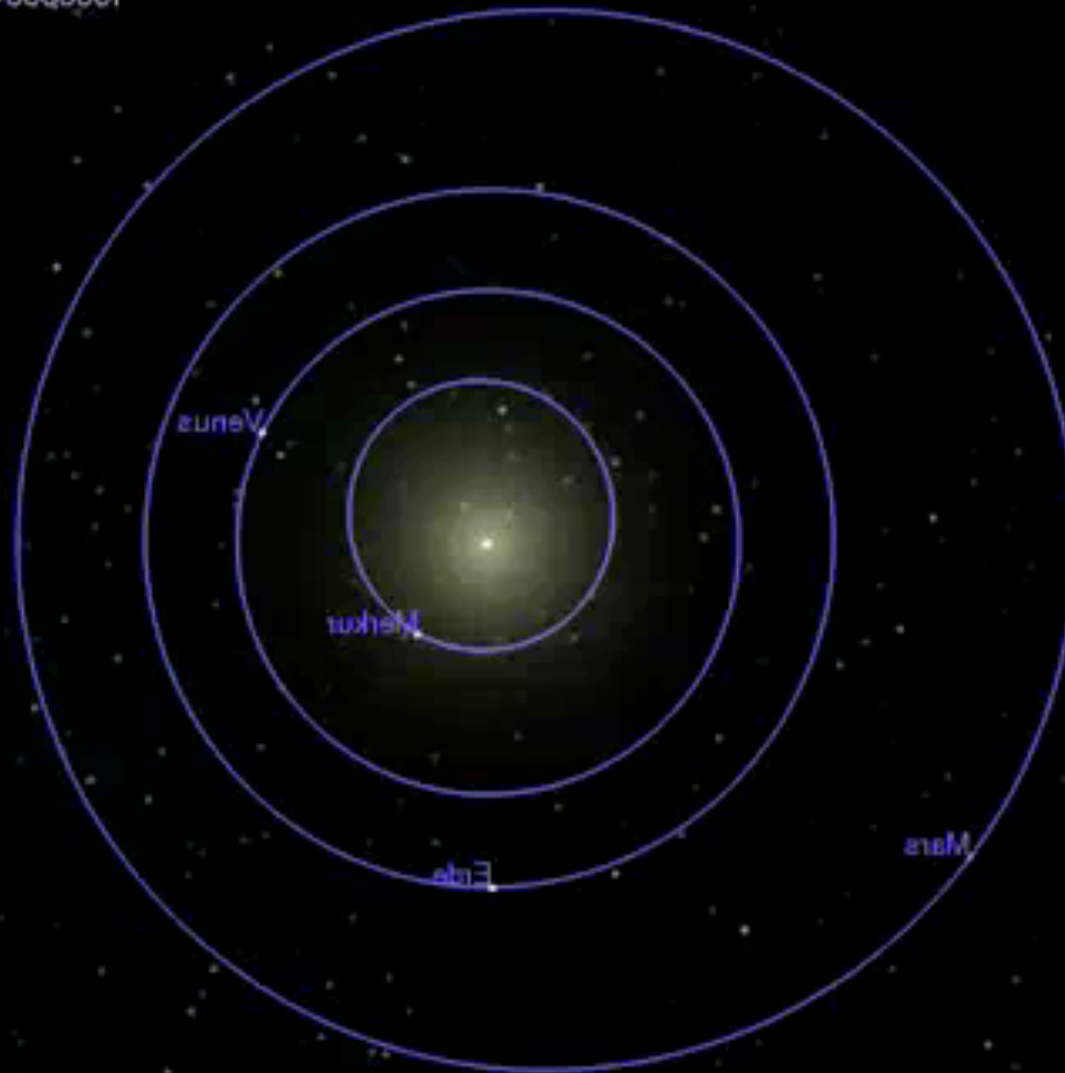
- Cleans rock surfaces with a brush
- Places and holds the APXS and MAHLI instruments
- Acquires samples of rock or soil with a powdering drill or scoop
- Sieves the samples (to 150 μm or 1 mm) and delivers them to instruments or an observation tray
- Exchanges spare drill bits



Curiosity's Sampling System

How do we get there?

2011 Nov 26 12:21:59 STD
1000000x schneller (Vergleichen)

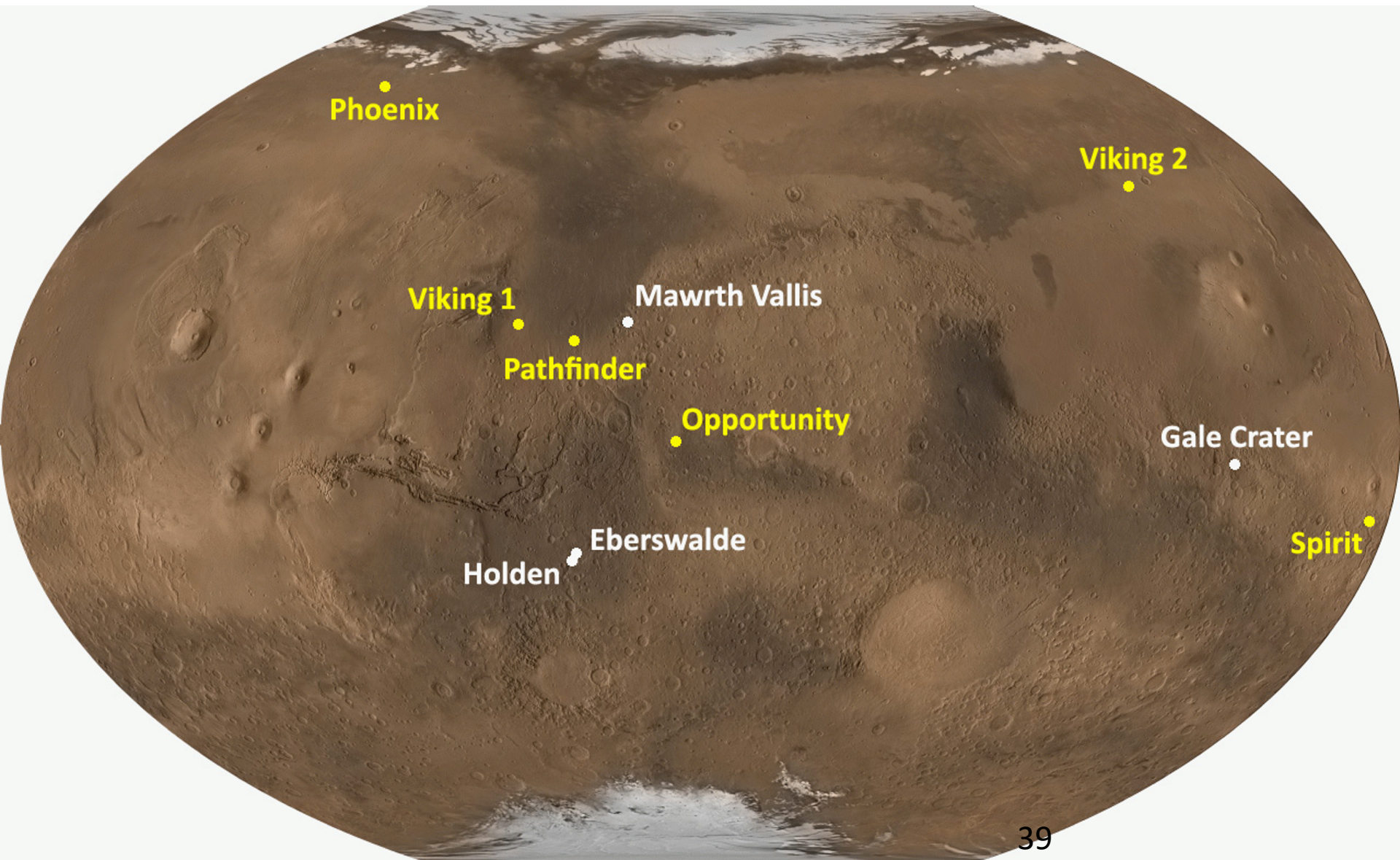


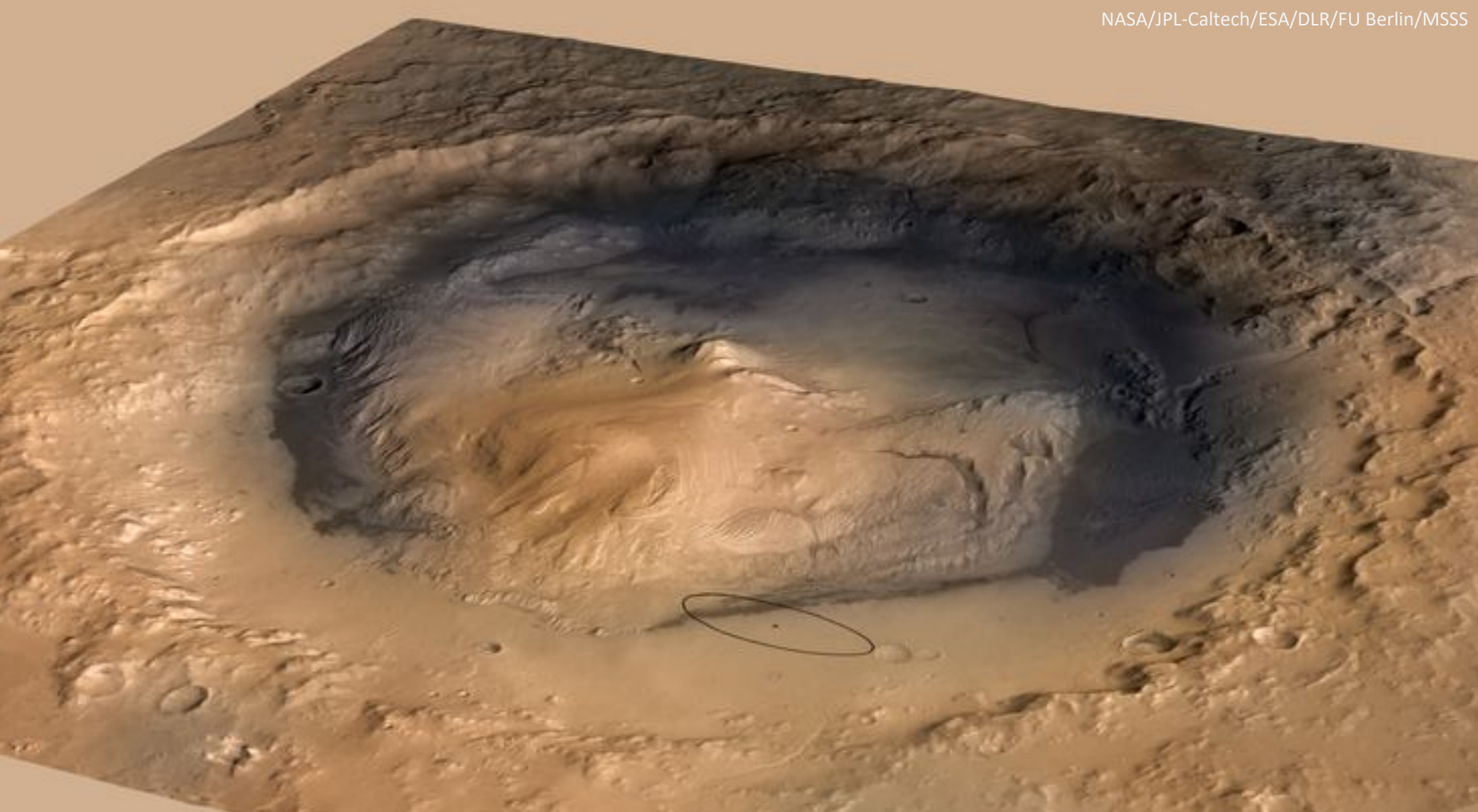
FOV: 18° 11' 19.3" (1.00x)

Geschwindigkeit: 0.00000 m/s

Mars Landing Sites

(Previous Missions and MSL Final Candidates)





150-km Gale Crater contains a 5-km high mound of stratified rock. Strata in the lower section of the mound vary in mineralogy and texture, suggesting that they may have recorded environmental changes over time.





NASA/JPL-Caltech/MSSS



Kicking up dust just prior to landing



**August 5, 2012: “Touchdown confirmed.”
“Time to see where our Curiosity will take us.”**



Curiosity self-portrait at Rocknest

Assembled from 55
MAHLI images

Shows four scoop
trenches and wheel
scuff



30'

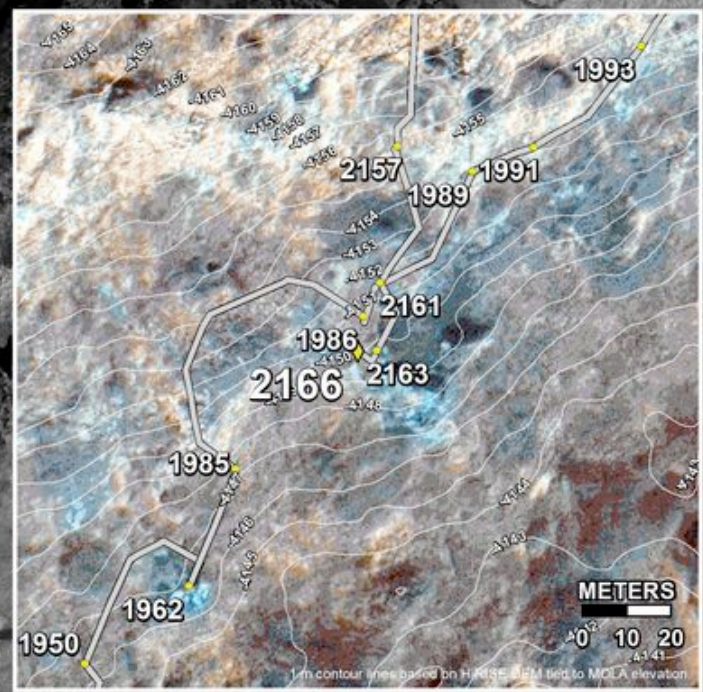
- 4°36'S

-4°40'S

-4°44'S



- **Rover Way Points** = Traverse

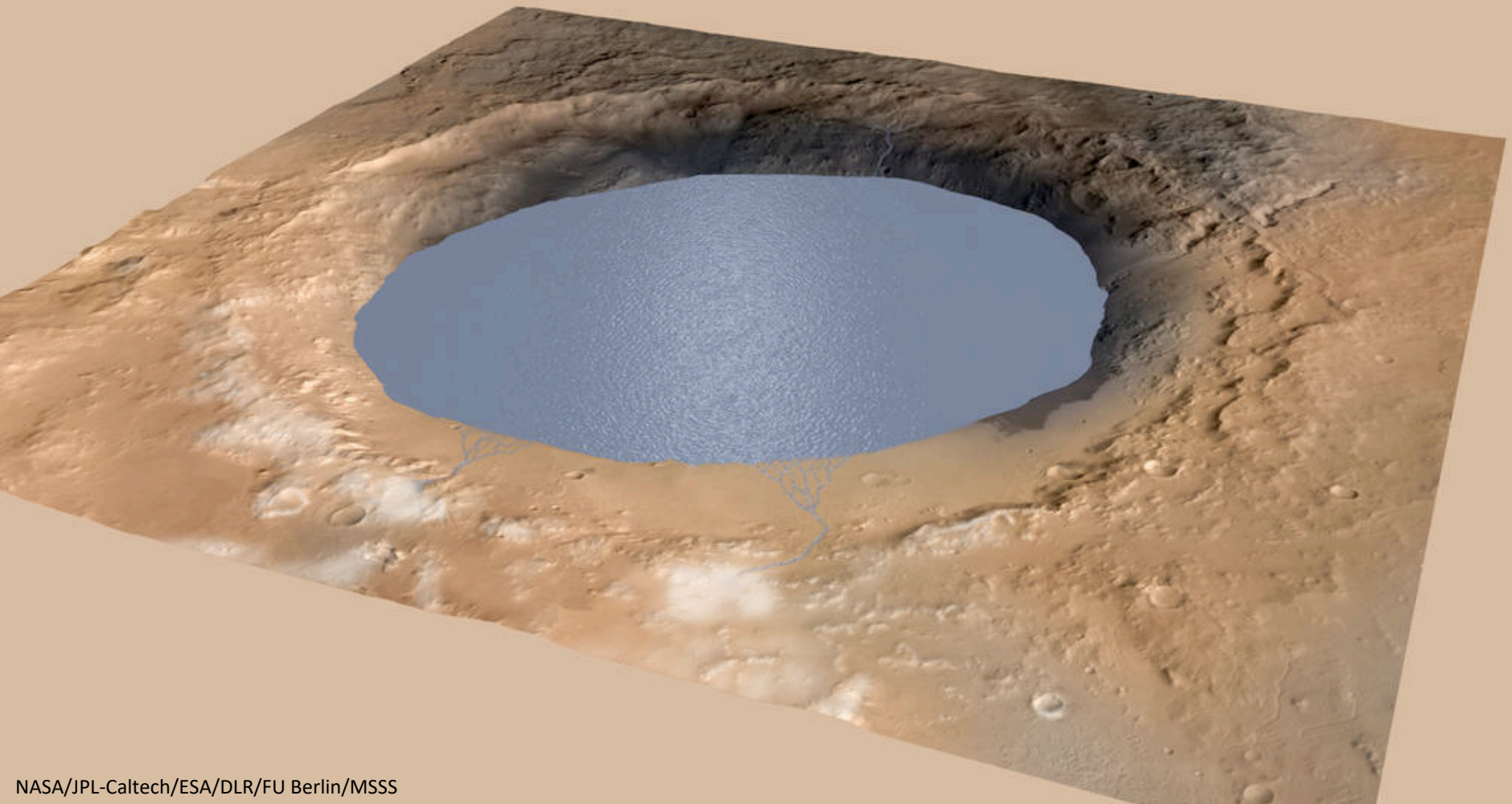


An Ancient Habitable Environment at Yellowknife Bay

- **The regional geology and fine-grained rock suggest that the John Klein site was at the end of an ancient river system or within an intermittently wet lake bed**
- **The mineralogy indicates sustained interaction with liquid water that was not too acidic or alkaline, and low salinity. Furthermore, conditions were not strongly oxidizing.**
- **Key chemical ingredients for life are present, such as carbon, hydrogen, nitrogen, oxygen, phosphorus, and sulfur**
- **The presence of minerals in various states of oxidation would provide a source of energy for primitive organisms**

Notes on Organics

- This is the first-ever detection of a martian organic chemical.
- It took many analyses of rocks and soils, as well as additional analyses of blanks and calibration standards on Mars and on Earth, in order to verify this discovery.
- SAM detected simple hydrocarbon molecules in which some of the hydrogen was replaced with chlorine. This could have happened on Mars, or within the instrument, through reaction with perchlorate compounds that are known to be widespread on Mars.
- Simple organic molecules do not require biology for their formation. However, they are building blocks of life. More importantly, we now can study what environments preserve organics on Mars' surface, increasing our ability to search for other life-related materials.



NASA/JPL-Caltech/ESA/DLR/FU Berlin/MSSS

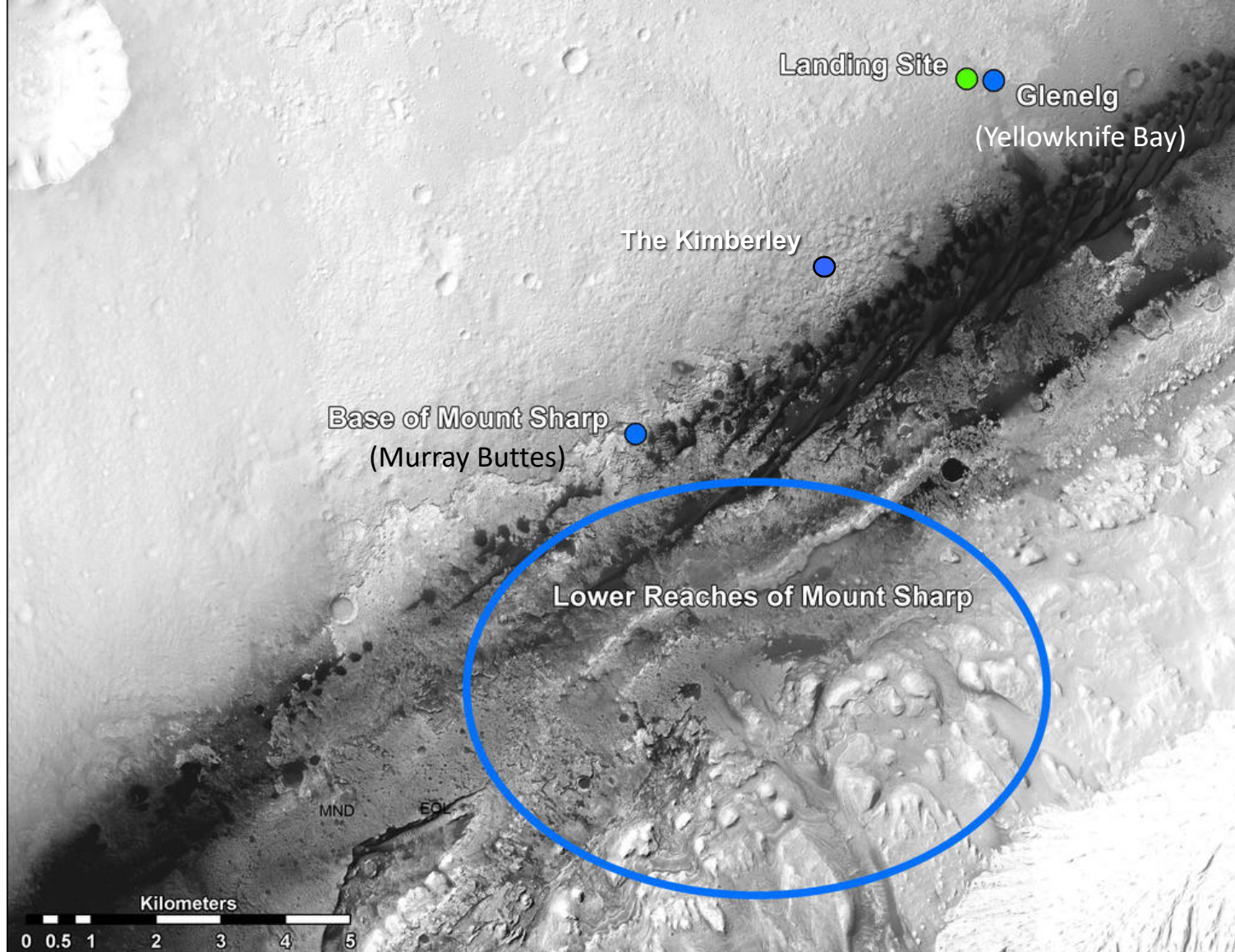


An illustration of lake partially filling Gale Crater. If such a lake existed for millions of years, it would have required a more humid climate and active hydrological cycle.

Recent Accomplishments

Since our January Science Team Meeting:

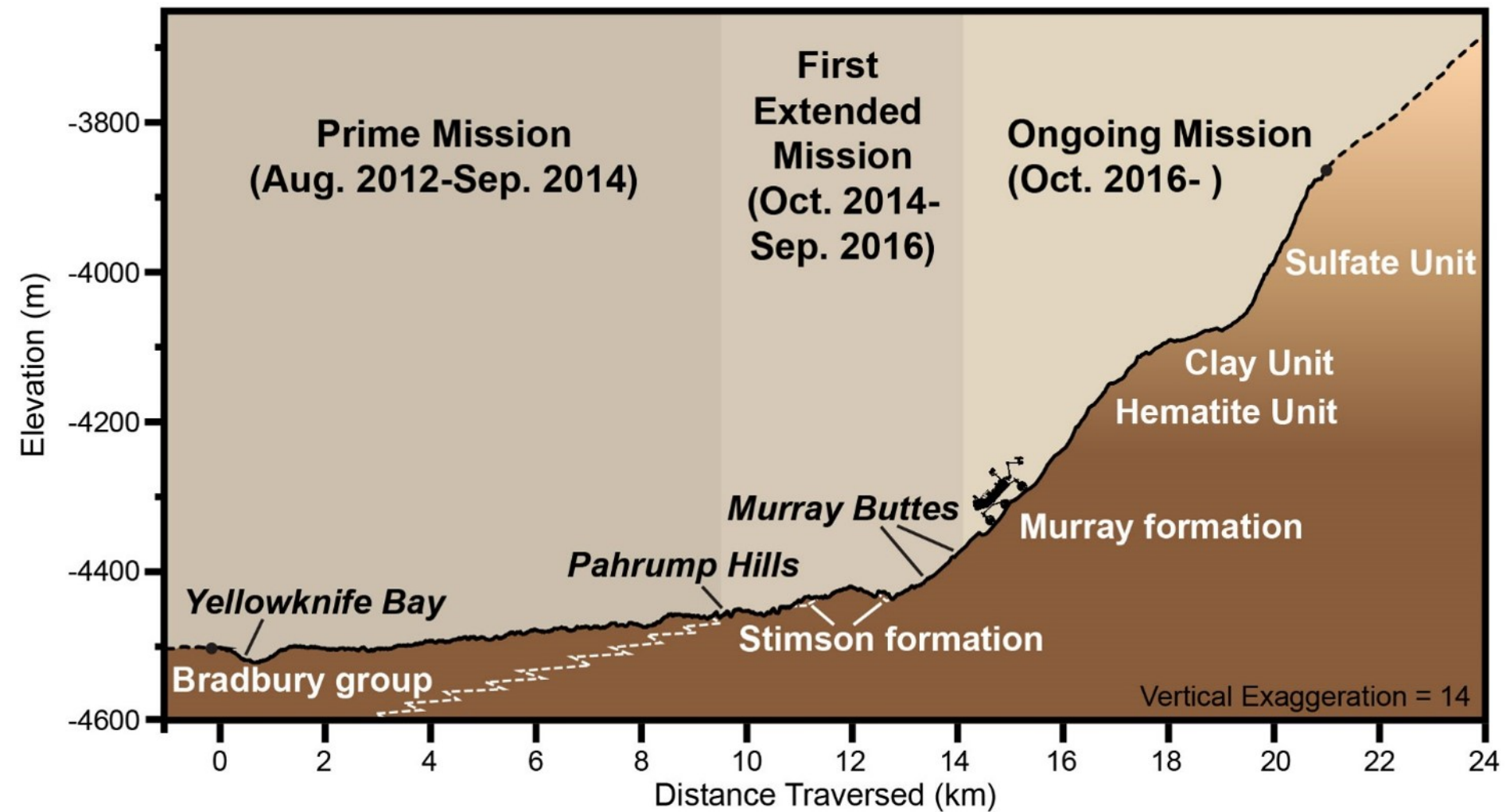
- Completed six Earth years and three Mars years of surface operations
- Returned to drilling!
 - Non-percussive drilling attempts at Lake Orcadie (1 and 2)
 - Successful percussive drilling and sample delivery at Duluth
 - Percussive drilling attempts at Ailsa Craig, Voyageurs, and Inverness
 - Successful percussive drilling and sample delivery at Stoer
- Completion of the Vera Rubin Ridge campaign minus sampling
- Development of science goals and strategic route for the clay-bearing unit
- Execution of a global dust storm campaign
- Publication of organics and methane *Science* papers and press conference
- Publication of Bagnold Dunes Phase 2 special issue in *GRL*
- Publication of 2017 solar particle event special issue in *Space Weather*



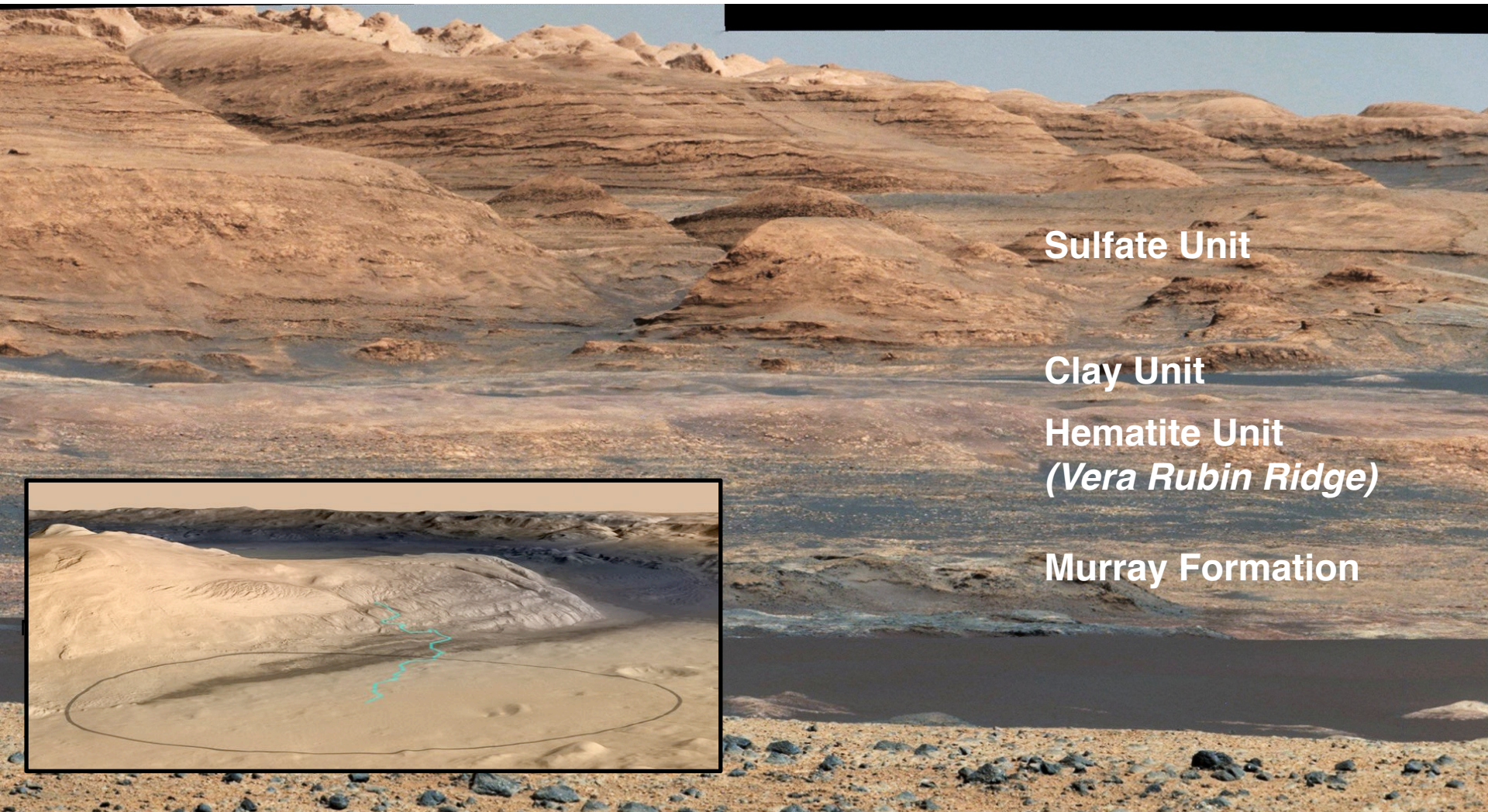
NASA/JPL-Caltech/Univ. of Arizona



Curiosity's ultimate goal is to explore the lower reaches of the 5-km high Mount Sharp



Curiosity is more steeply climbing Mount Sharp in the second two-year extension of its mission

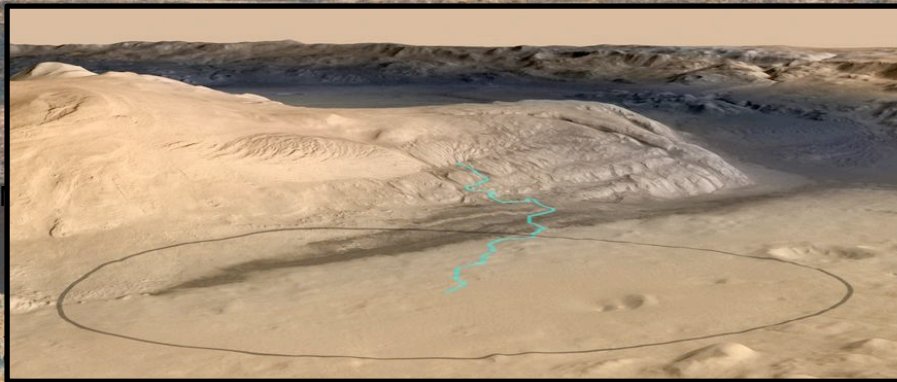


Sulfate Unit

Clay Unit

Hematite Unit
(*Vera Rubin Ridge*)

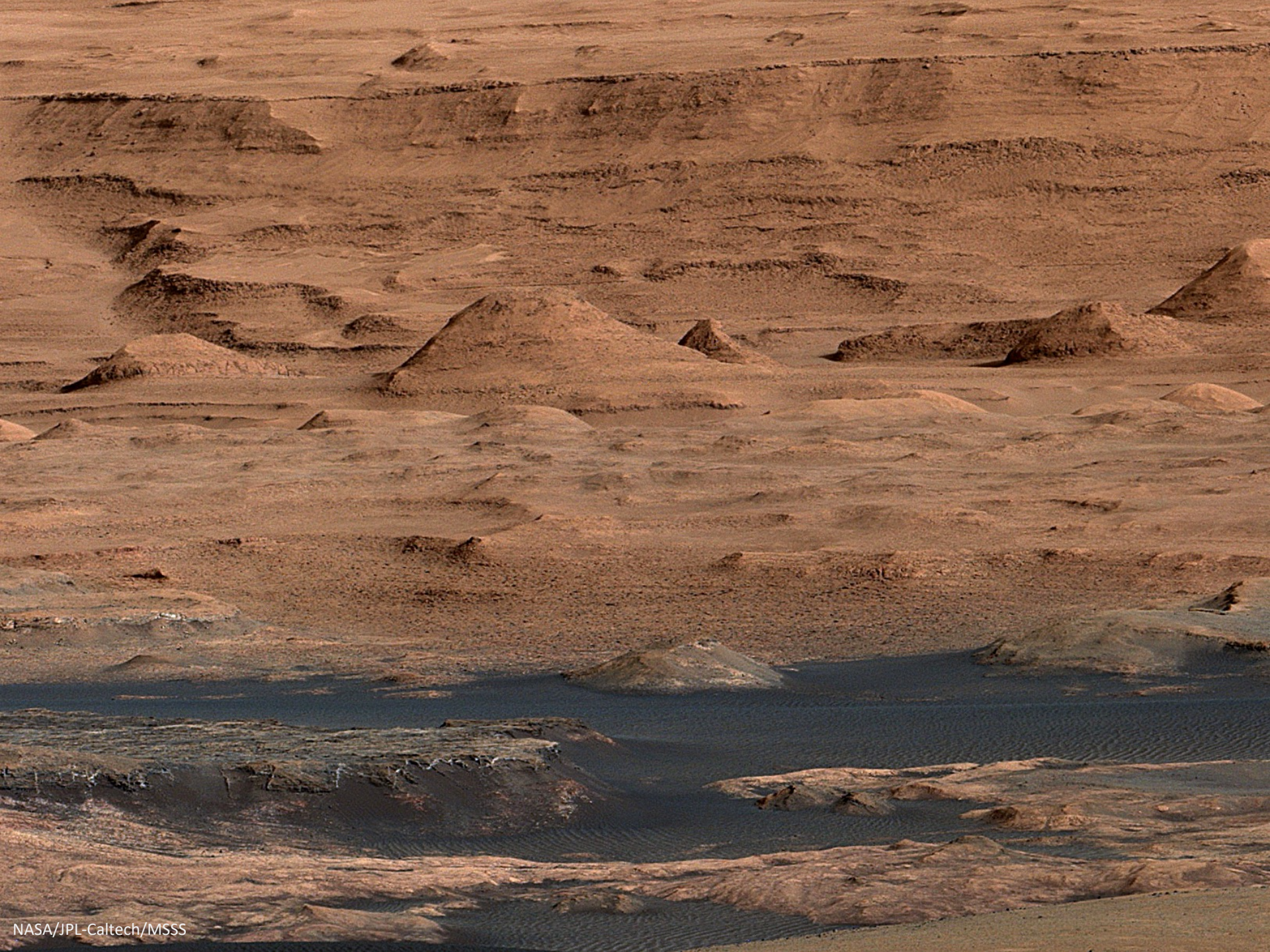
Murray Formation



NASA/JPL-Caltech/MSSS



Curiosity's Extended Mission will explore Mt. Sharp, with an emphasis on understanding the subset of habitable environments that preserve organic carbon

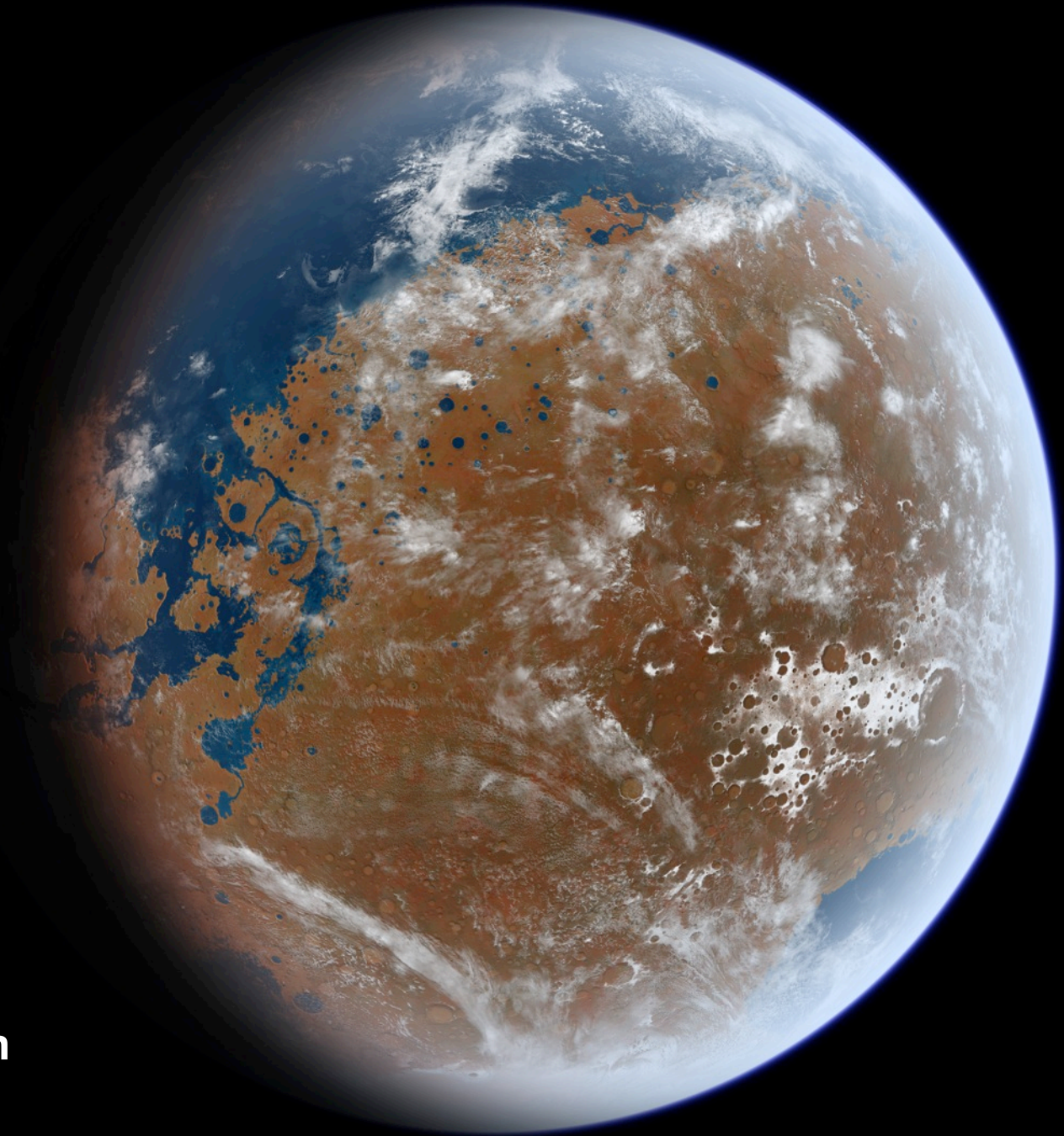


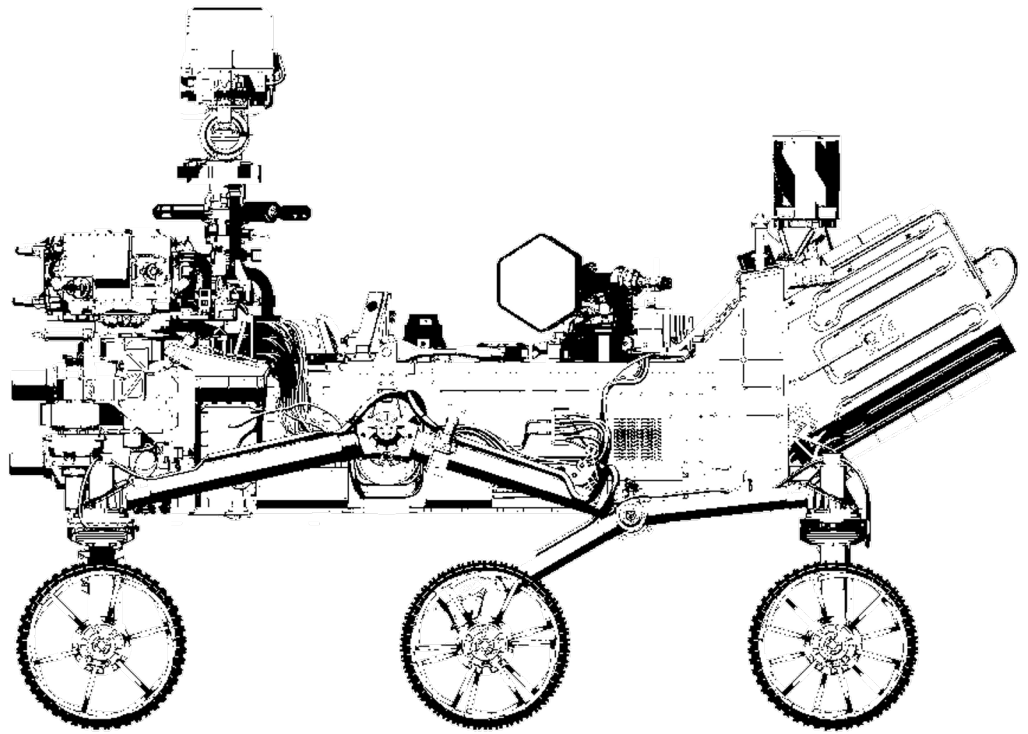


Mars exploration results are strong evidence that early Mars was warm and wet.

Aqueous minerals, high silica deposits, long-lived lakes in Gale Crater would be possible only with an active hydrological cycle and other bodies of water.

Mars was habitable, could have been inhabited, may be today, and could be in the future.



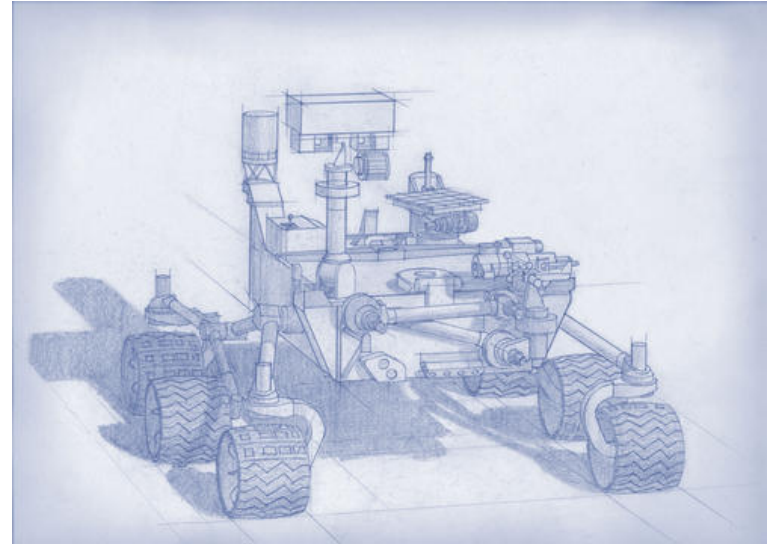


MARS 2020

Project Overview

Salient Features

- *Category: 1*
- *Risk Class: A-tailored*
- *Directed, JPL in-house implementation*
- *High heritage MSL design*
- *Modifications only as necessary to accommodate new payload and Sampling / Caching System (SCS)*
- *Planetary Protection Category V per Program direction*



Science (<https://mars.nasa.gov/mars2020/mission/science/objectives/>)

- *Characterize the processes that formed and modified the geologic record within a field exploration area on Mars selected for evidence of an astrobiologically-relevant ancient environment and geologic diversity.*
- *Perform the following astrobiologically relevant investigations on the geologic materials at the landing site:*
 1. *Determine the habitability of an ancient environment.*
 2. *For ancient environments interpreted to have been habitable, search for materials with high biosignature preservation potential.*
 3. *Search for potential evidence of past life using the observations regarding habitability and preservation as a guide.*
- *Assemble rigorously documented and returnable cached samples for possible future return to Earth.*
- *Contribute to the preparation for human exploration of Mars by making significant progress towards filling at least one major Strategic Knowledge Gap (SKG).*

Mission Overview



LAUNCH

- Atlas V 541 vehicle
- Launch Readiness Date: July 2020
- Launch window: July/August 2020

CRUISE/APPROACH

- ~7 month cruise
- Arrive Feb 2021

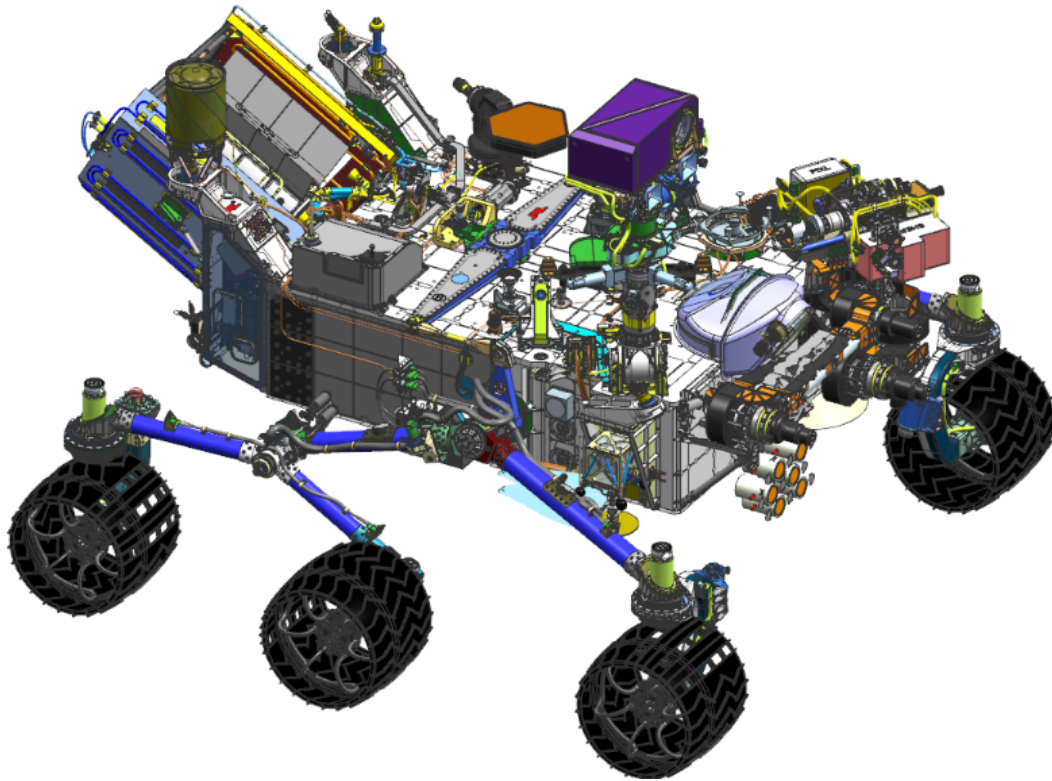
ENTRY, DESCENT & LANDING

- MSL EDL system (+ [Range Trigger](#) and [Terrain Relative Navigation](#)): guided entry and powered descent/Sky Crane
- 16 x 14 km landing ellipse (range trigger baselined)
- Access to landing sites $\pm 30^\circ$ latitude, ≤ -0.5 km elevation
- Curiosity-class Rover

SURFACE MISSION

- 20 km traverse distance capability
- [Enhanced surface productivity](#)
- [Qualified to 1.5 Martian year lifetime](#)
- Seeking signs of past life
- Returnable cache of samples
- Prepare for human exploration of Mars

Mars 2020 Rover Concept



Stays the Same as MSL

- Avionics
- Power
- GN&C
- Telecom
- Thermal
- Mobility

Changed

- New Science Instrument Suite
- New Sampling Caching System
- Modified Chassis
- Modified Rover Harness
- Modified Surface FSW
- Modified Rover Motor Controller
- Modified Wheels

Mars 2020 Rover

